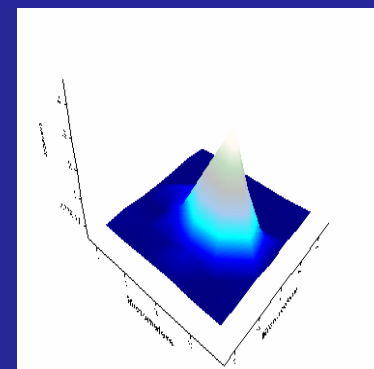
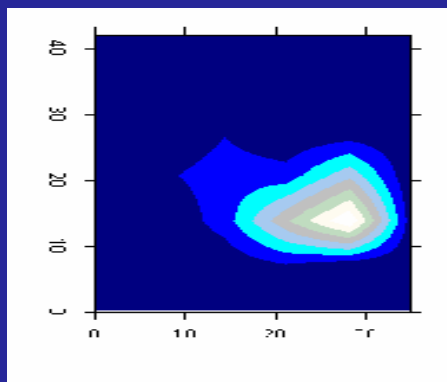
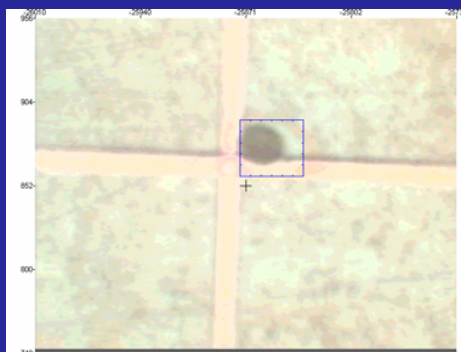
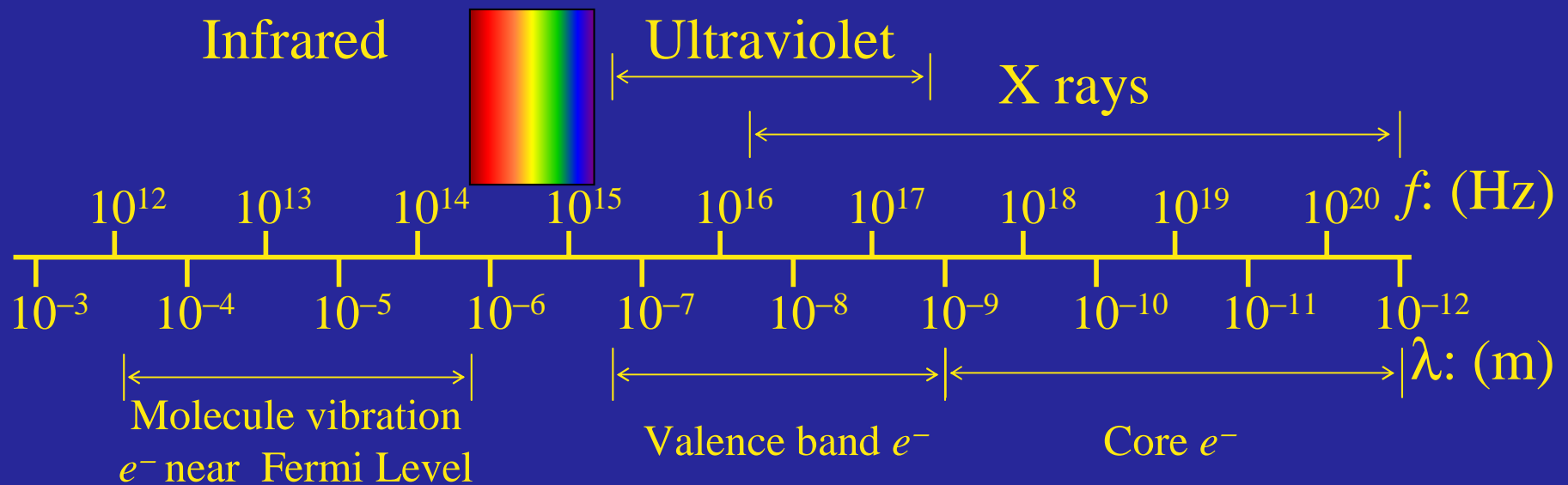


# Frontiers of **infrared** spectroscopy: Infrared Beamlines and Applications in Biology, Geology and Environmental Remediation

*Carol Hirschmugl*

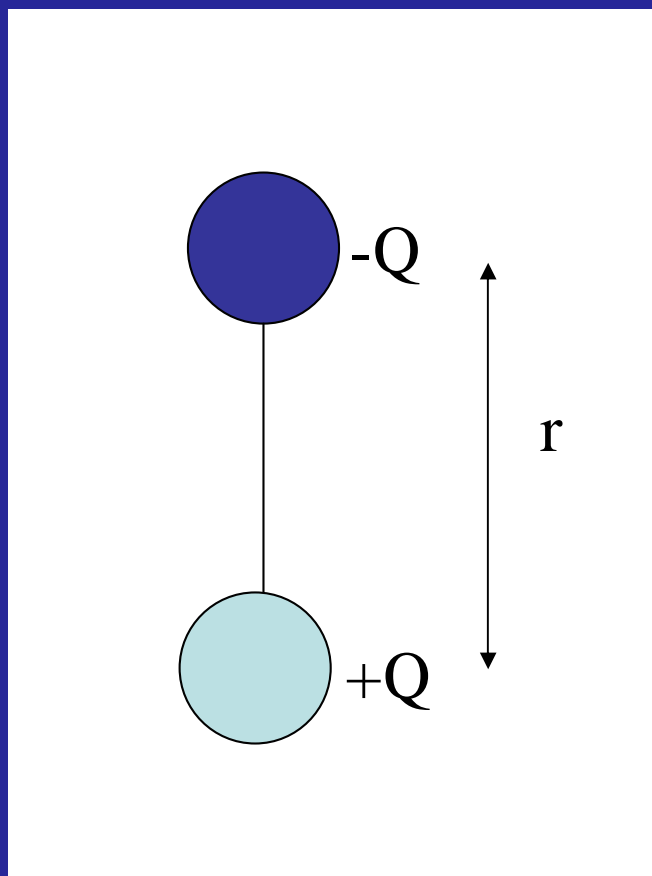


# What is Infrared Spectroscopy?

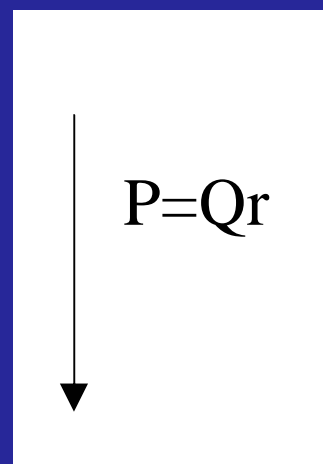


Absorption of Low Energy (IR) Photons causes:  
Displacement of atoms (Vibrations)  
Functional groups (CH, OH, etc.) within larger molecules  
exhibit similar absorption energies

# What is a Static Dipole



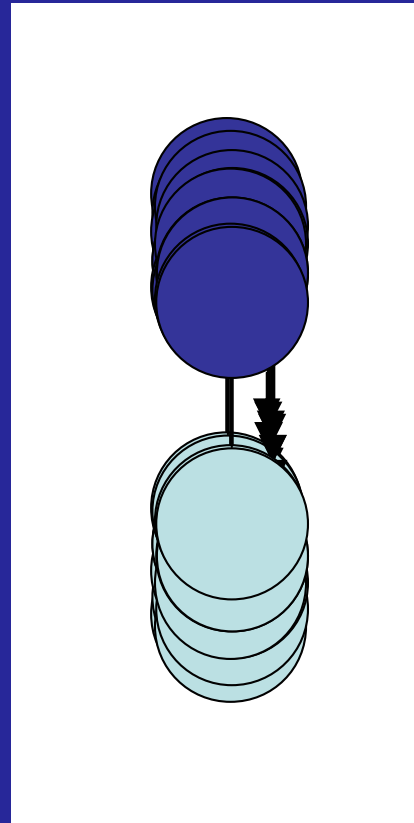
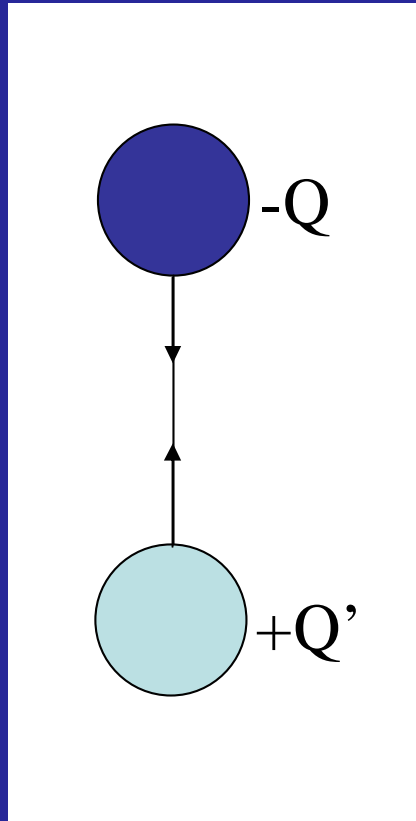
$Q$  is charge  
 $r$  is distance between the charges  
 $P$  is dipole



Which of the following molecules are dipoles:

$O_2$ ,  $CO$ ,  $N_2$ ,  $NO$ ?

# What is a Dynamic Dipole



Separated charge distribution that is moving

Which of the following would be Dynamic Dipoles:

Vibrating  $N_2$

Vibrating  $CO$

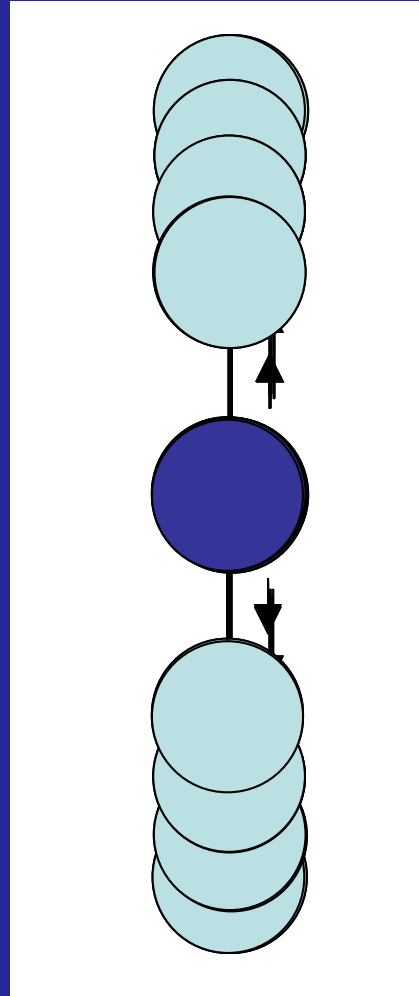
Vibrating  $O_2$

Vibrating  $NO$

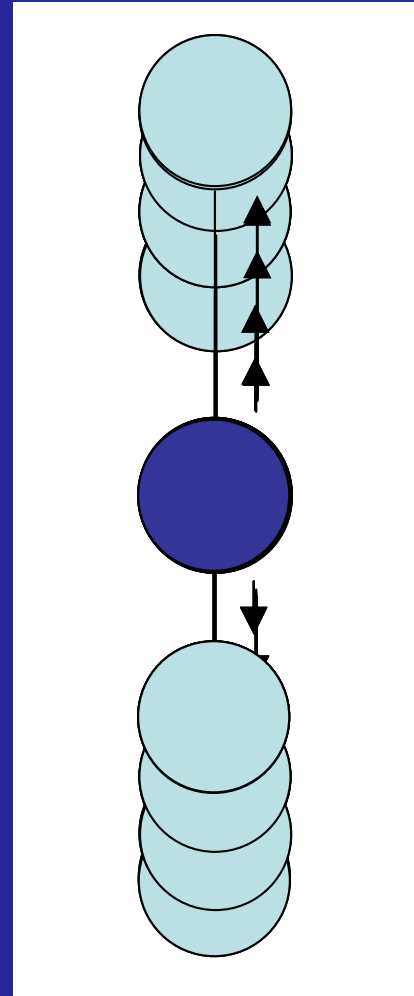
Vibrating  $CO_2$ ????

Ionic bonds in molecules can have infrared active vibrations,  
Covalent bonds in molecules are not infrared active

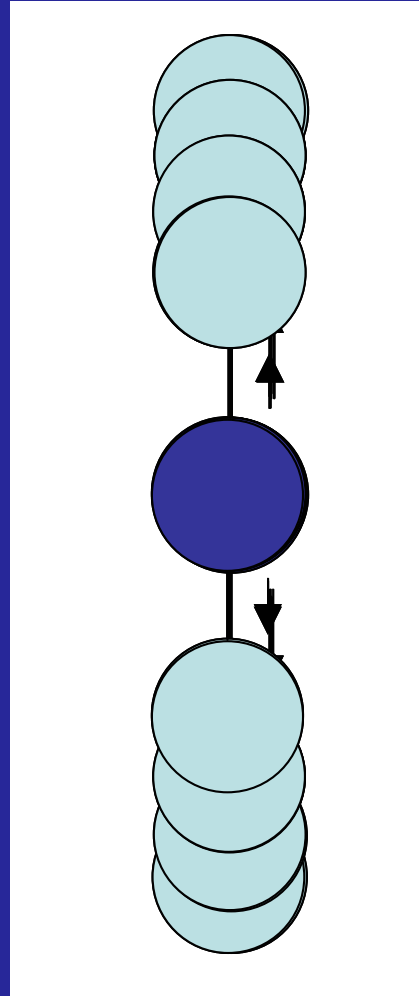
# $\text{CO}_2$ Vibrations: Stretching Modes



# $\text{CO}_2$ Vibrations: Stretching Modes

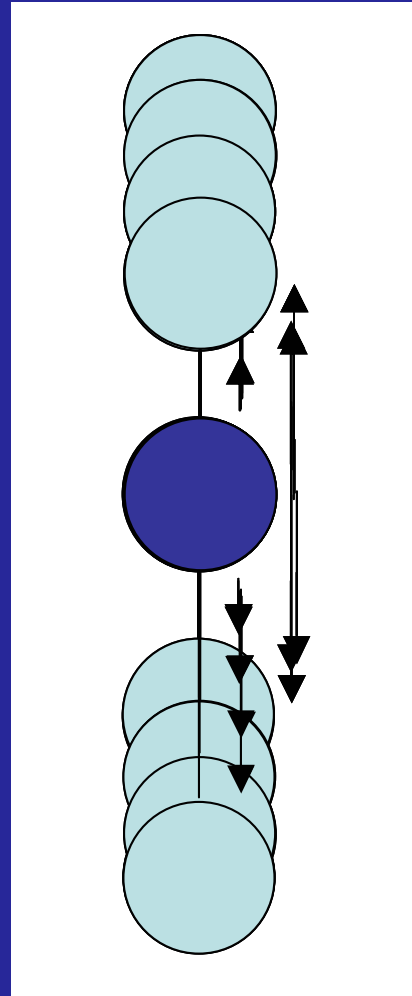


# CO<sub>2</sub> Vibrations: Net Dipoles

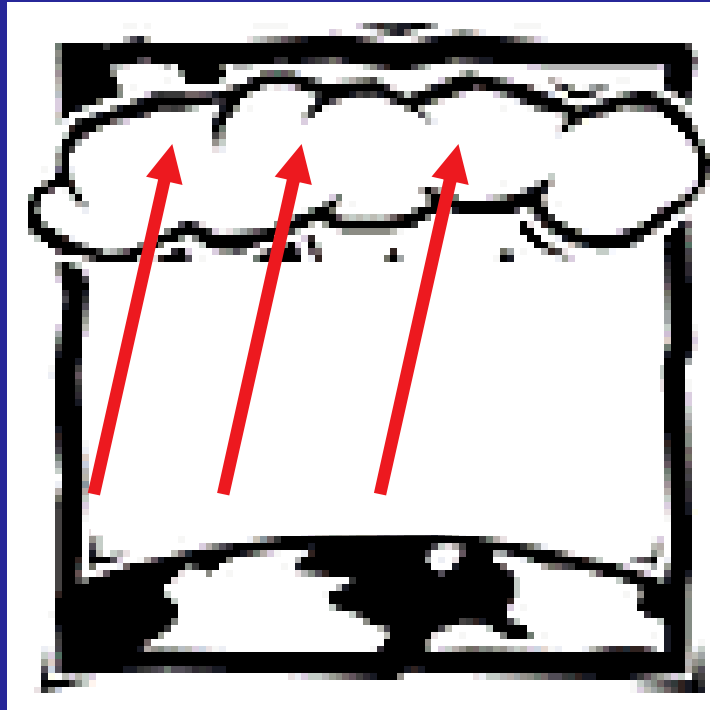




# CO<sub>2</sub> Vibrations: Net Dipoles



# What are greenhouse gases?

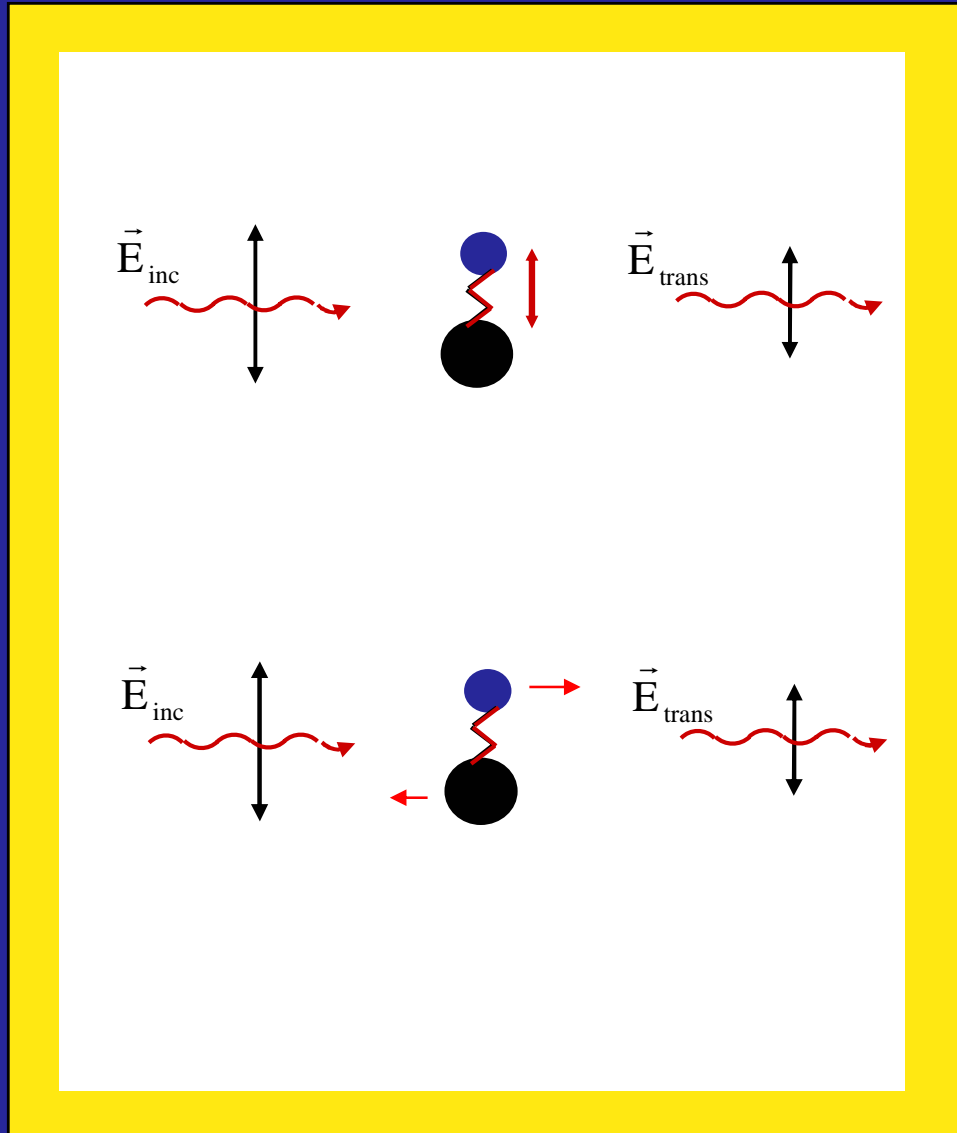


Air:

$N_2, O_2, CO_2, CH_4, H_2O$

Which are absorbing the infrared radiation reflecting from the earth?

# Vibrational Strengths



High Frequency Vibrations:  
Stretching Modes

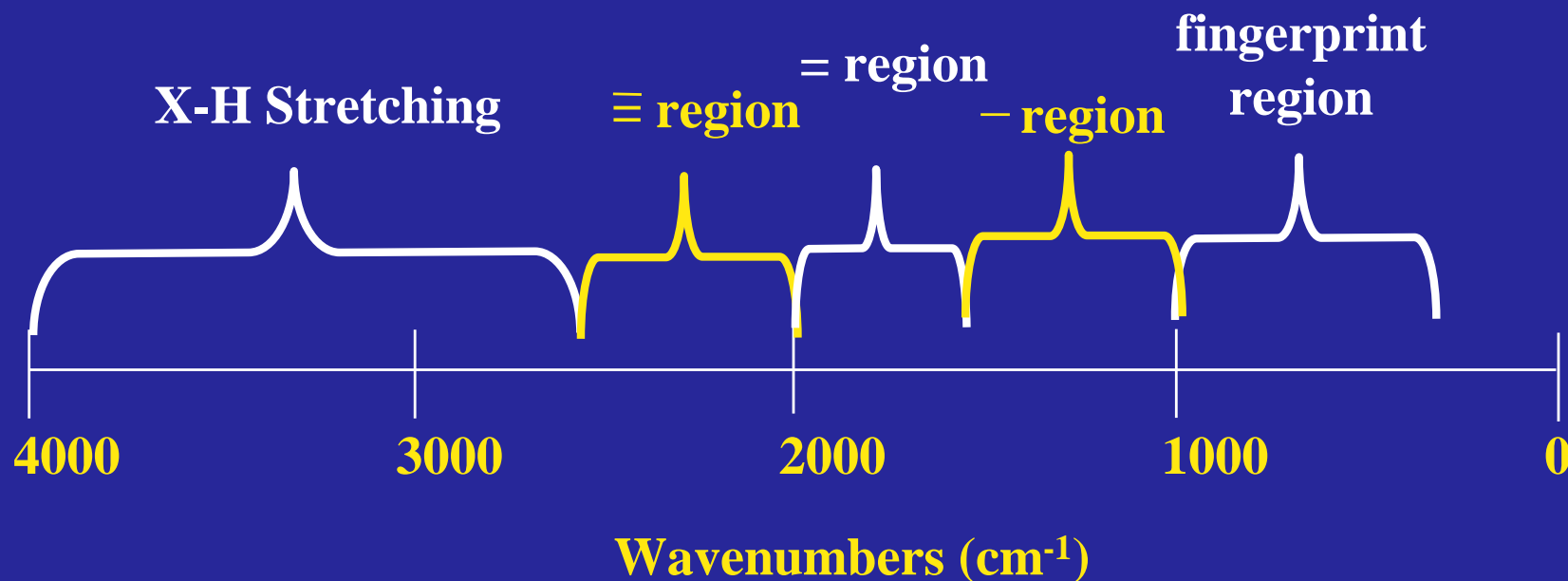
Lower Frequency Vibrations:  
Rotational Modes (bending and  
wagging modes)

# Vibrational Strengths

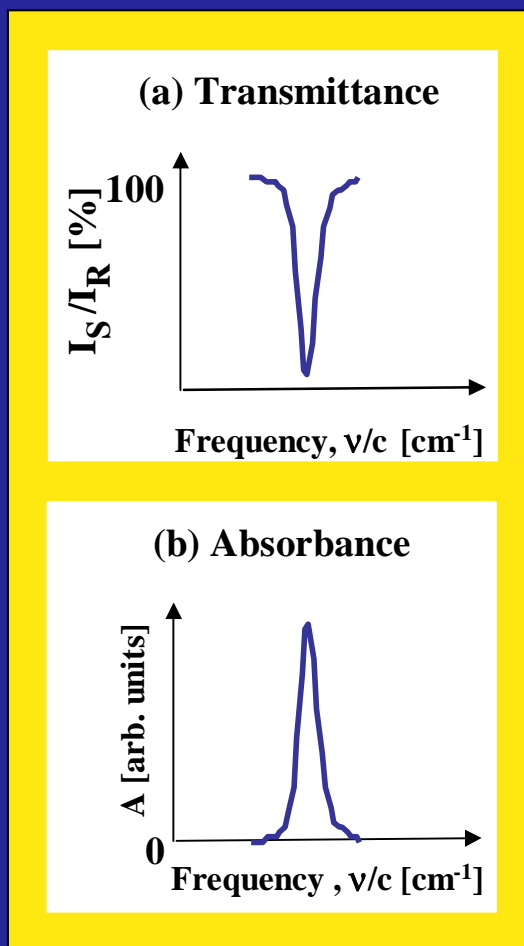


$$\nu = \frac{1}{2\pi} \sqrt{\frac{k}{m}}$$

Atoms in Typical Functional Groups include  
C, N, O, H, P, S, Si, Al



# Signatures in Infrared Spectroscopy?



The energy for each vibration is dependent upon the mass and spring constant, IR absorption spectroscopy is chemically specific.

# Sources of Infrared Light

## Blackbody Radiation

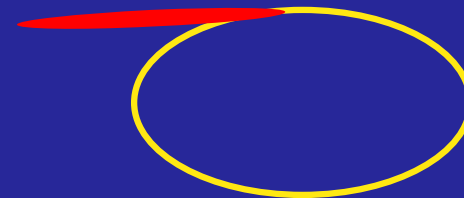


Hot filament or rod  
(Global or Nernst Glowar)

Flux radiated in all directions

Flux  $\propto$  Temperature

## Storage Ring Radiation

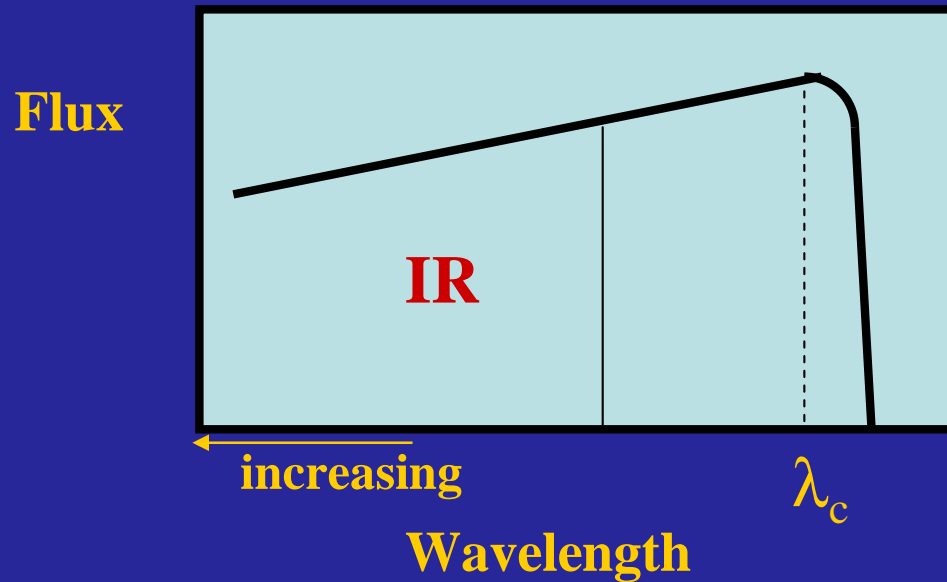


From relativistically accelerated  
charged particles

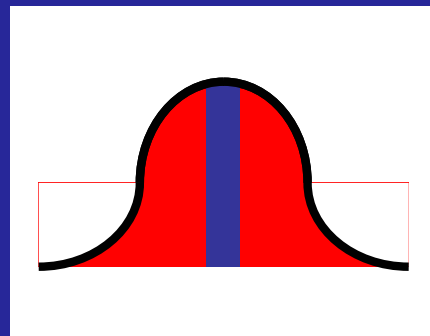
Flux radiated in a tangential,  
well defined cone

Flux  $\propto$  Beam current

# Synchrotron Characteristics: IRSR



**IR Flux  $\propto$  Beam Current**

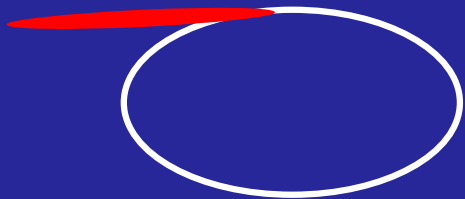


**Beam Profile**

**IR Source: wide beam profile**  
**Large opening angle**  
**Mixed polarization**

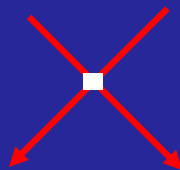
# Why use a storage ring source?

Storage Ring



Small source size  
Small solid angle

Spectromicroscopy



Very small sample  
 $1\ \mu\text{m} \times 1\ \mu\text{m}$

Grazing incidence  
reflection-absorption  
spectroscopy

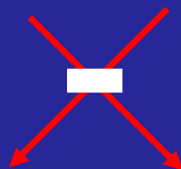


Very small  
angular acceptance  
 $< 5^\circ$

Globalar



Large source  
Large solid angle



Small sample  
 $30\ \mu\text{m} \times 30\ \mu\text{m}$



Small angular acceptance  
 $> 10^\circ$



# Advanced Light Source Infrared Facility

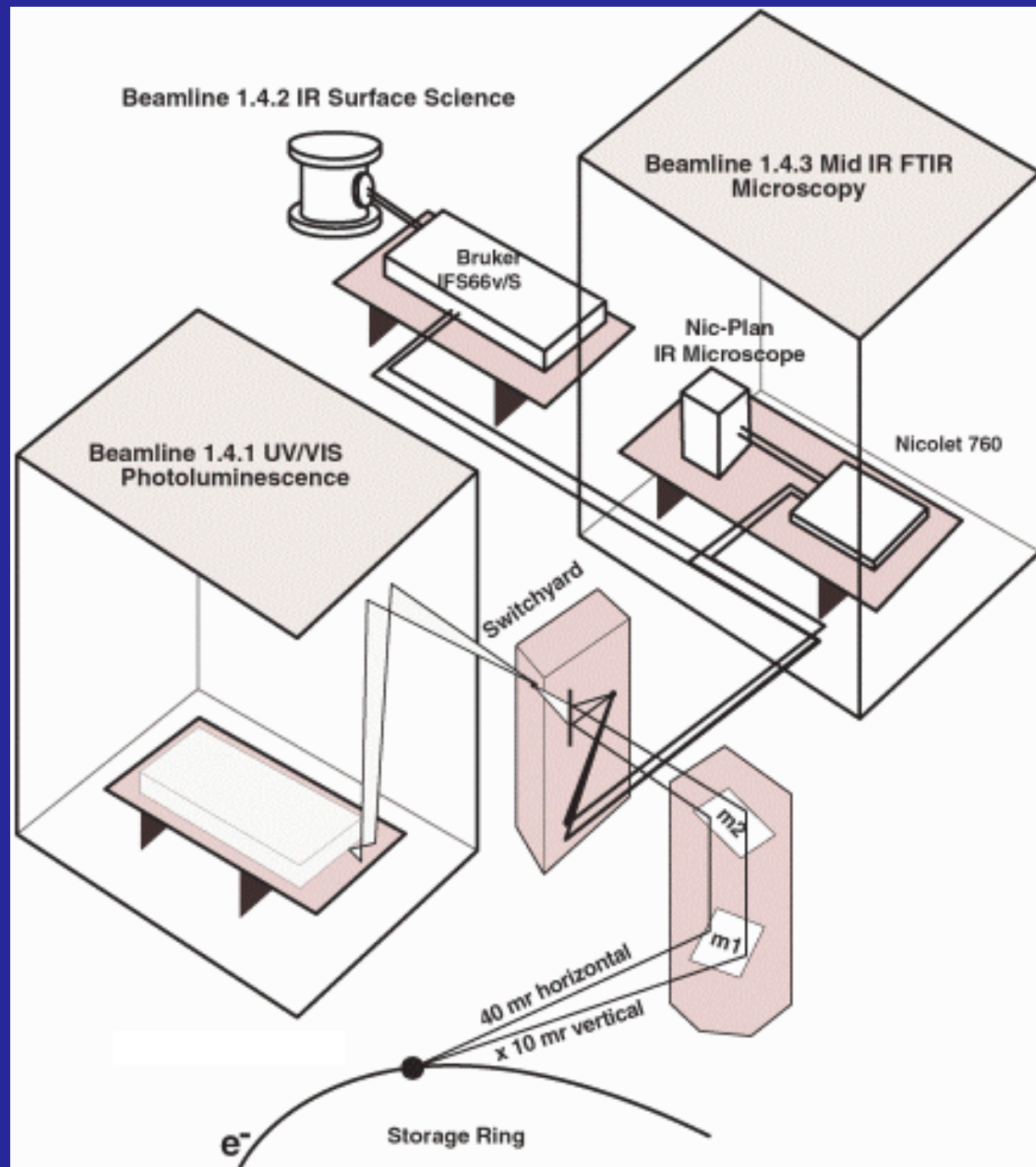
## Necessary Beamline Components:

### Components:

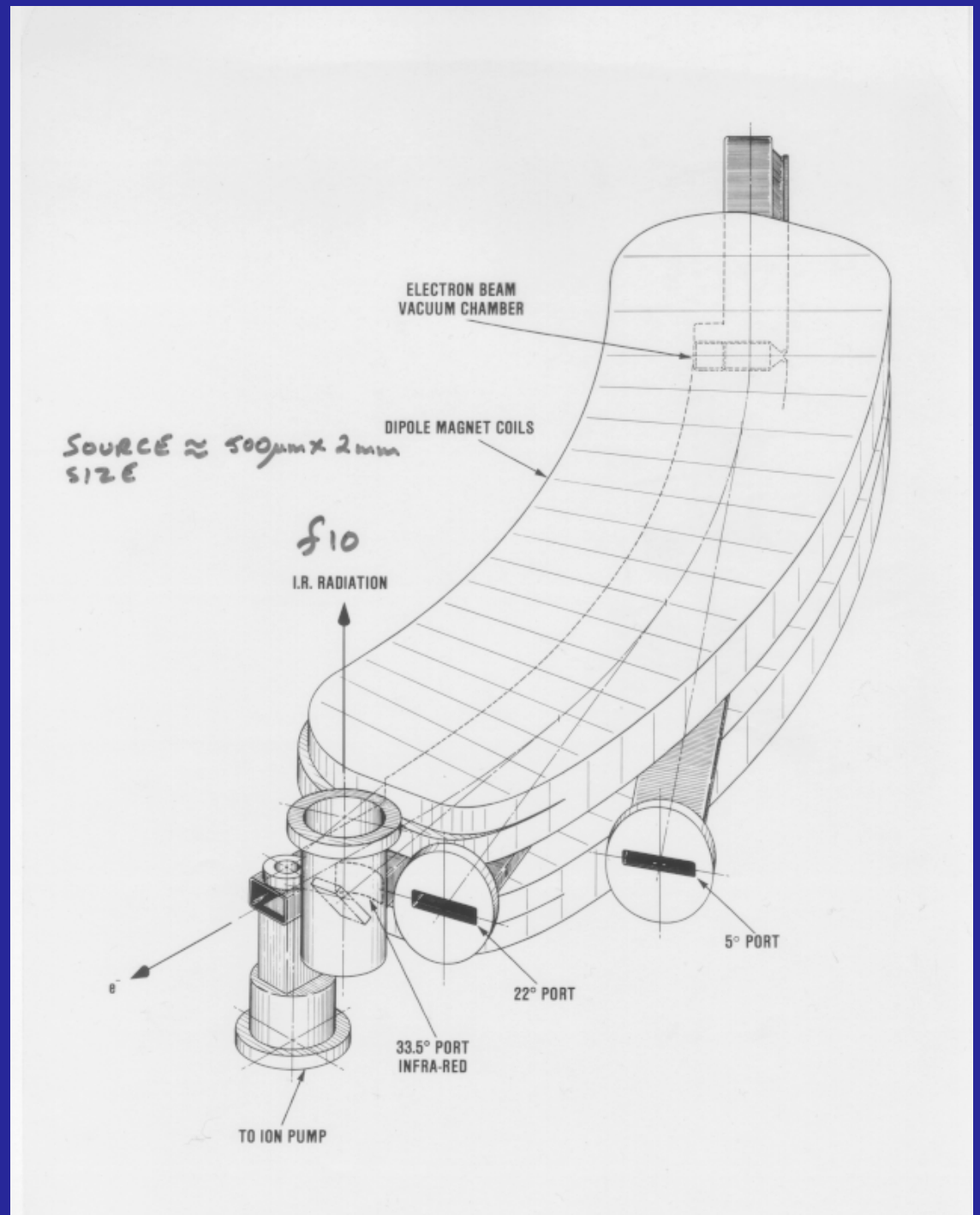
Large extraction port  
Large Water Cooled Flat Mirror  
Off-Axis Parabolic Mirror with long focal-lengths  
Wedged Diamond Window at focal point  
Collimating Optics  
Spectrometer: FTIR (Vacuum or Nitrogen Purge) w/ LN2 or LHe cooled detectors

## Optional Beamline Components: (Based on experiment)

Infrared Microscope  
Surface Science Chamber  
Beam position feedback system



# NSLS, BNL Vacuum Tank Schematic



NSLS, BNL  
Vacuum  
Tank  
IR and UV



CNR-439-85

NSLS, BNL  
Vacuum  
Tank  
IR and UV

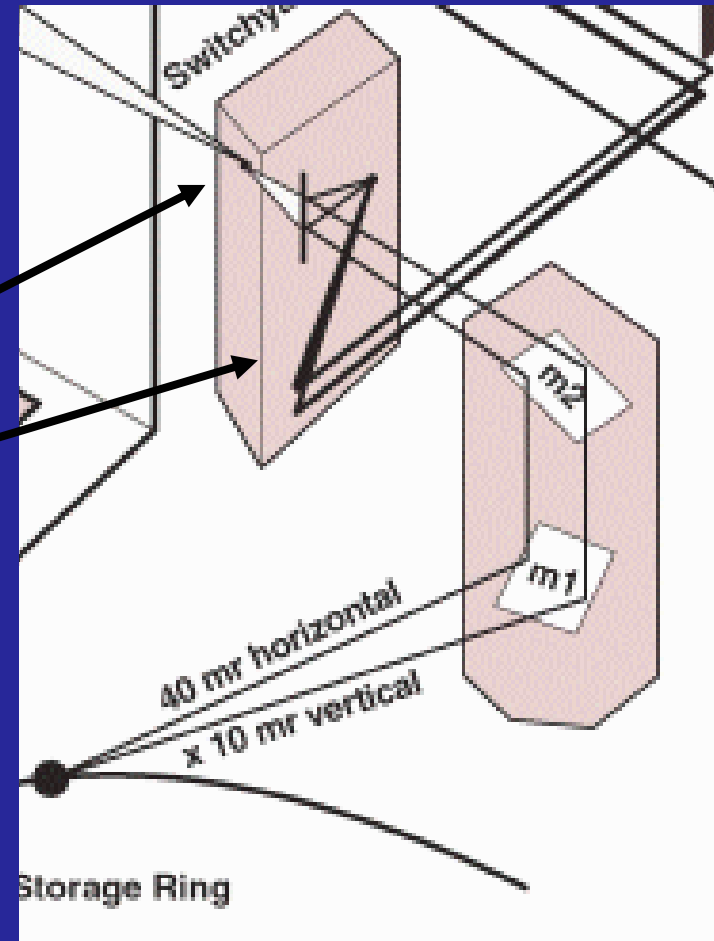
# Advanced Light Source Infrared Facility

m1: Large Water Cooled Flat Mirror

m2: Off-Axis Parabolic Mirror with long focal-lengths

Wedged Diamond Window at focal point

Collimating Optics



# Water Cooled Plane Mirror and Off Axis Paraboloid for ALS IR Beamline

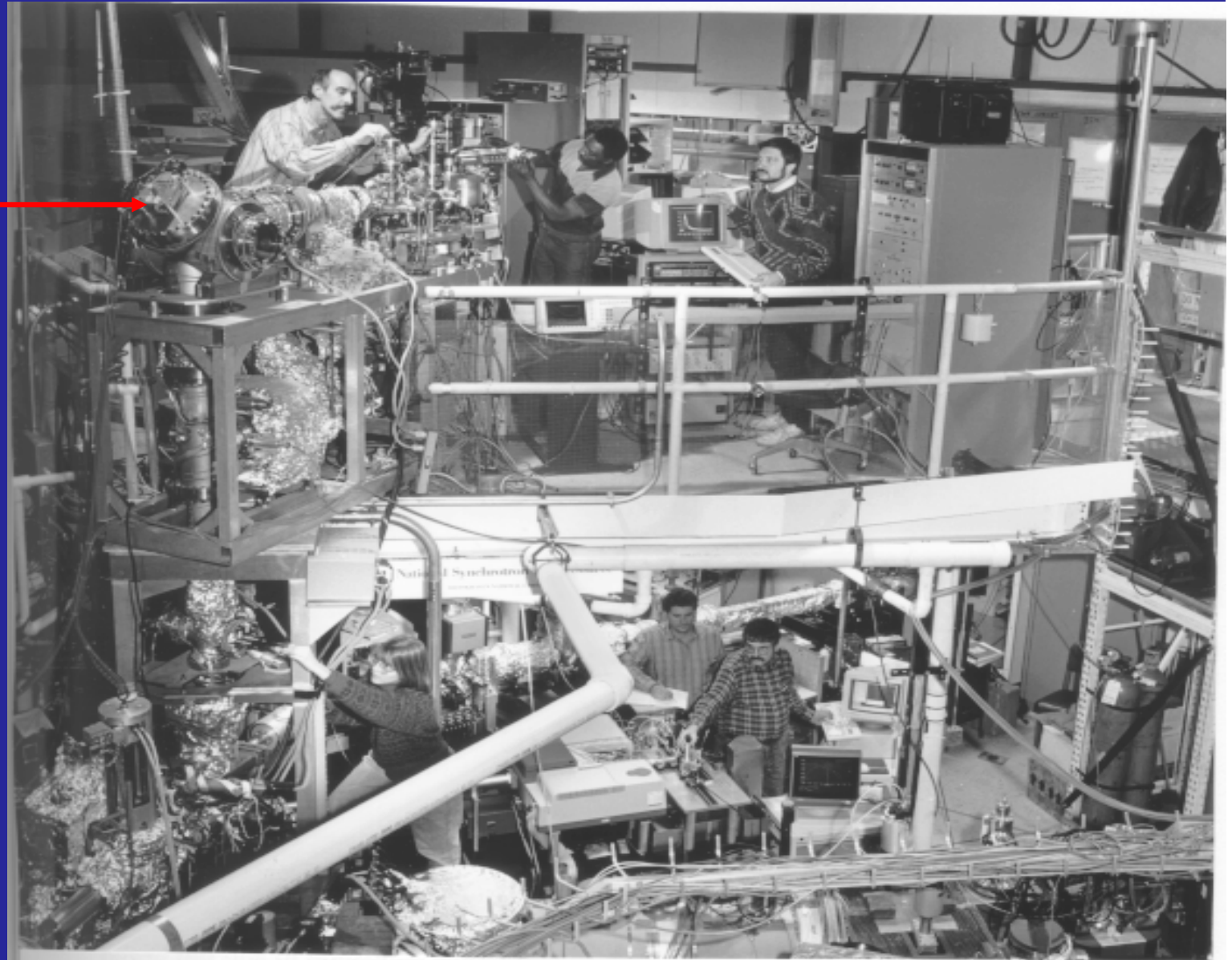


**Off Axis  
Paraboloid**

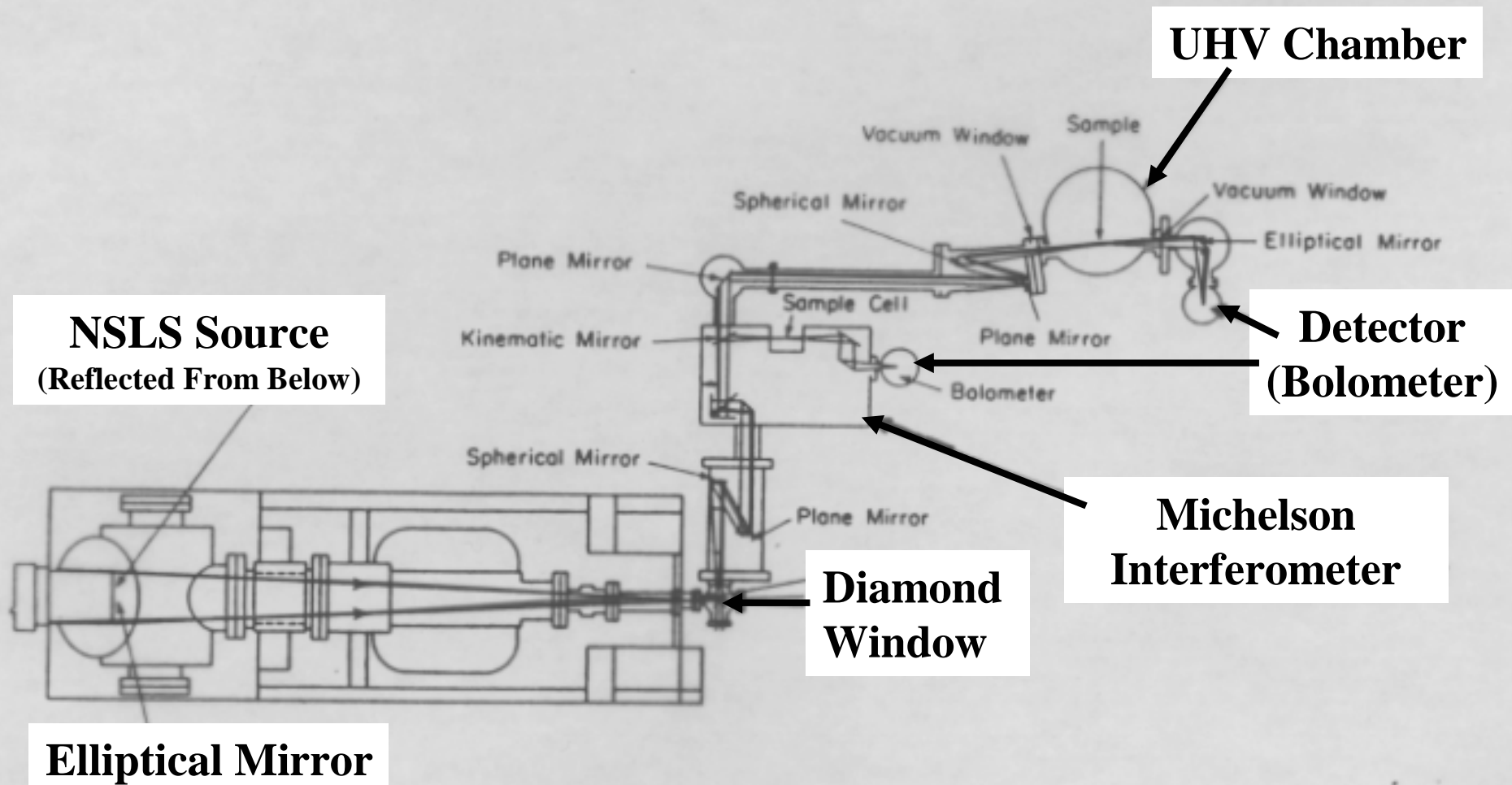
**Plane  
Mirror**

# NSLS U4IR and U2B Infrared Beamlines

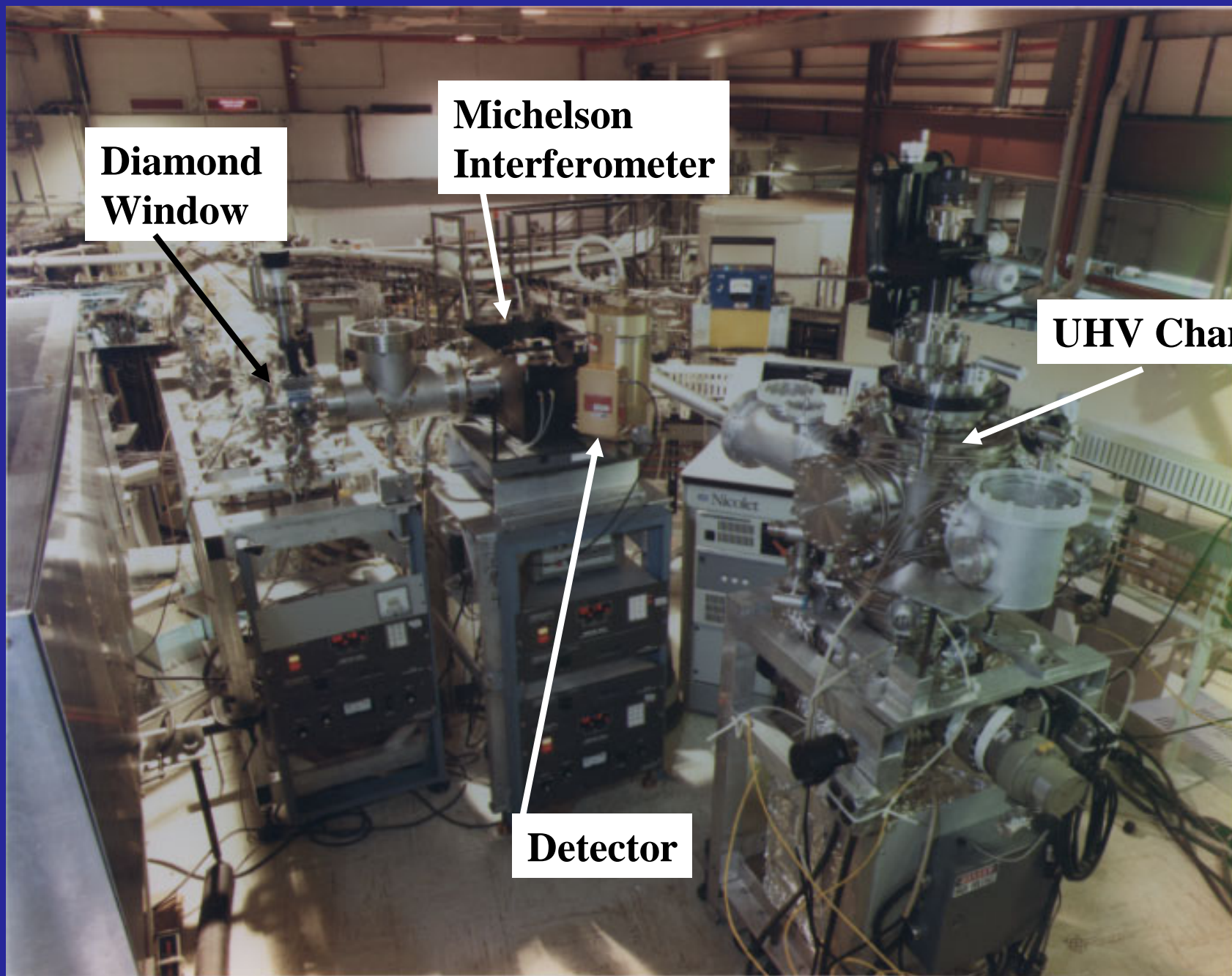
Off axis  
Paraboloid



# Infrared Beamline Overview: NSLS Beamline U4IR



# NSLS Beamline U4IR



**Diamond Window**

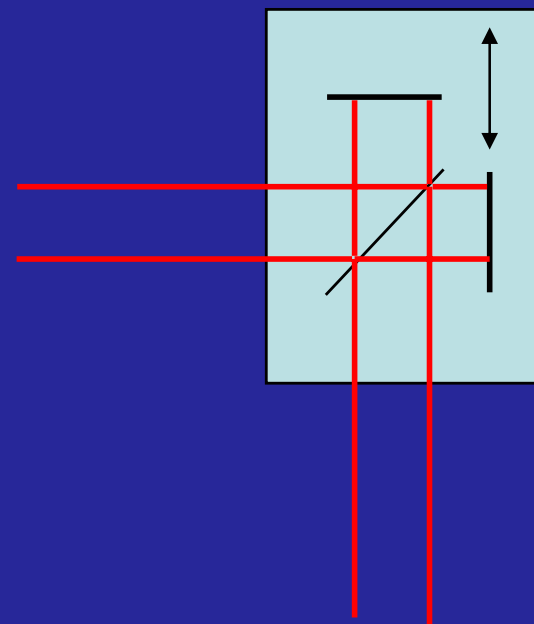
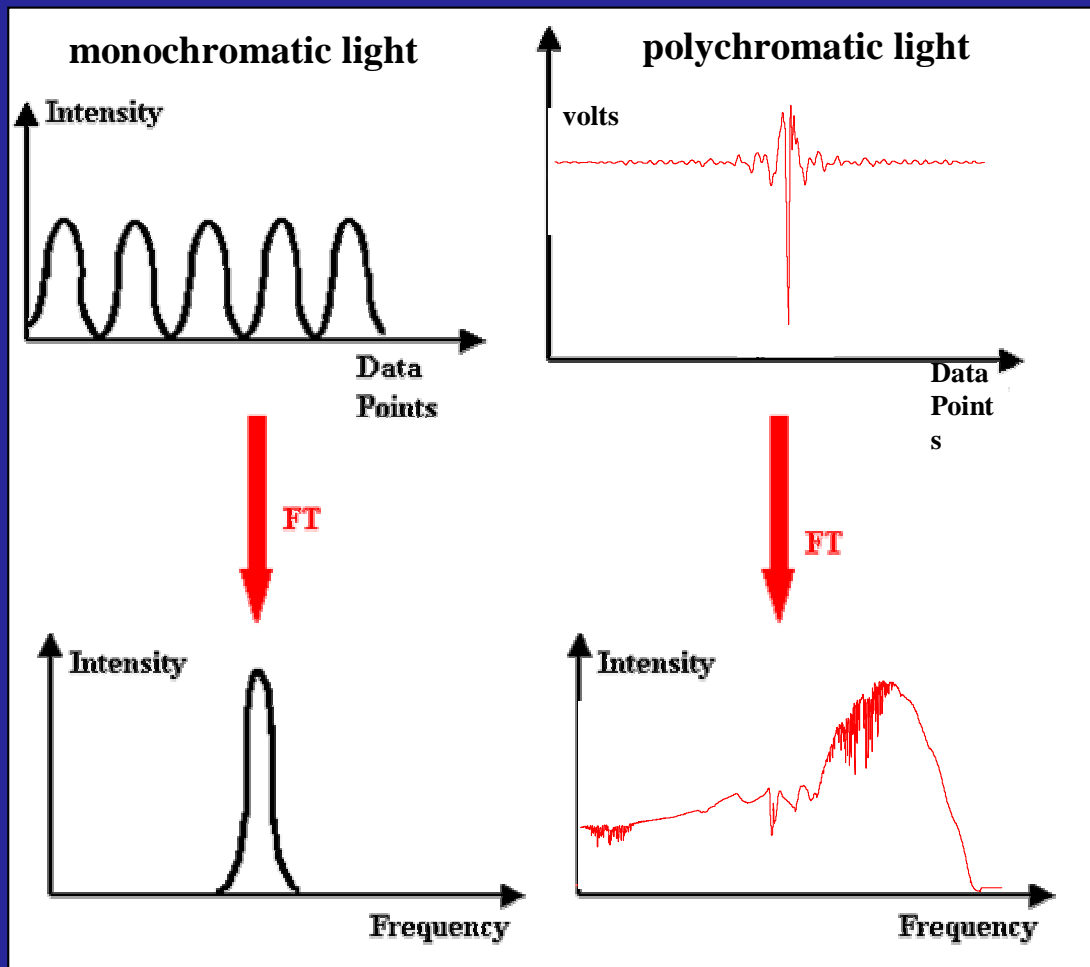
**Michelson Interferometer**

**UHV Chamber**

**Detector**



# FTIR Michelson Interferometer



# Rapid Scan Michelson Interferometer

**Broadband IR from  
Storage Ring  
( $1 \text{ cm}^{-1} - 40,000 \text{ cm}^{-1}$ )**

+

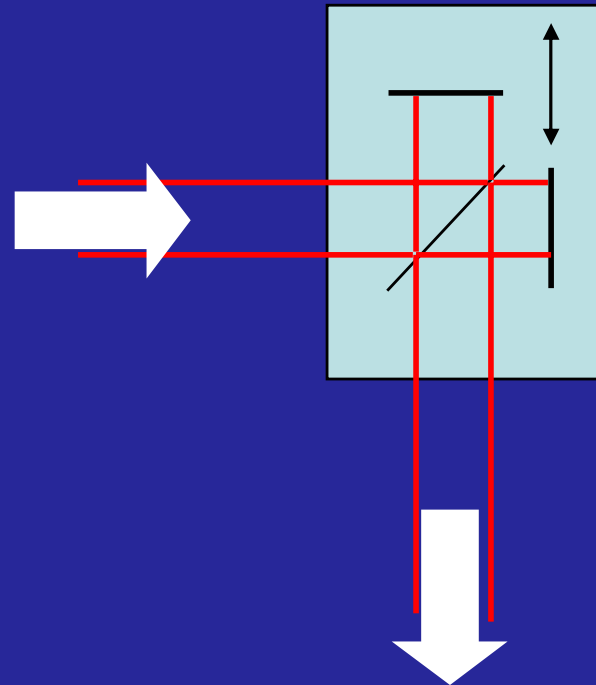
**“Noise” from  
Electron beam motion  
( e.g. 100 Hz)**

**Examples include:**

**RF Side bands  
(700-3000 Hz)**

**Ripple on Power Supplies  
(multiples of 60 hz)**

**etc.**



**Moving mirror  
velocity = 1 cm/s**

**Broadband IR Signal  
( $1 \text{ cm}^{-1} - 40,000 \text{ cm}^{-1}$ )**

+

**Signal Frequency “Noise”  
( $100 \text{ cm}^{-1}$  spike)**

# Thermo Nicolet Infrared Microscope at ALS

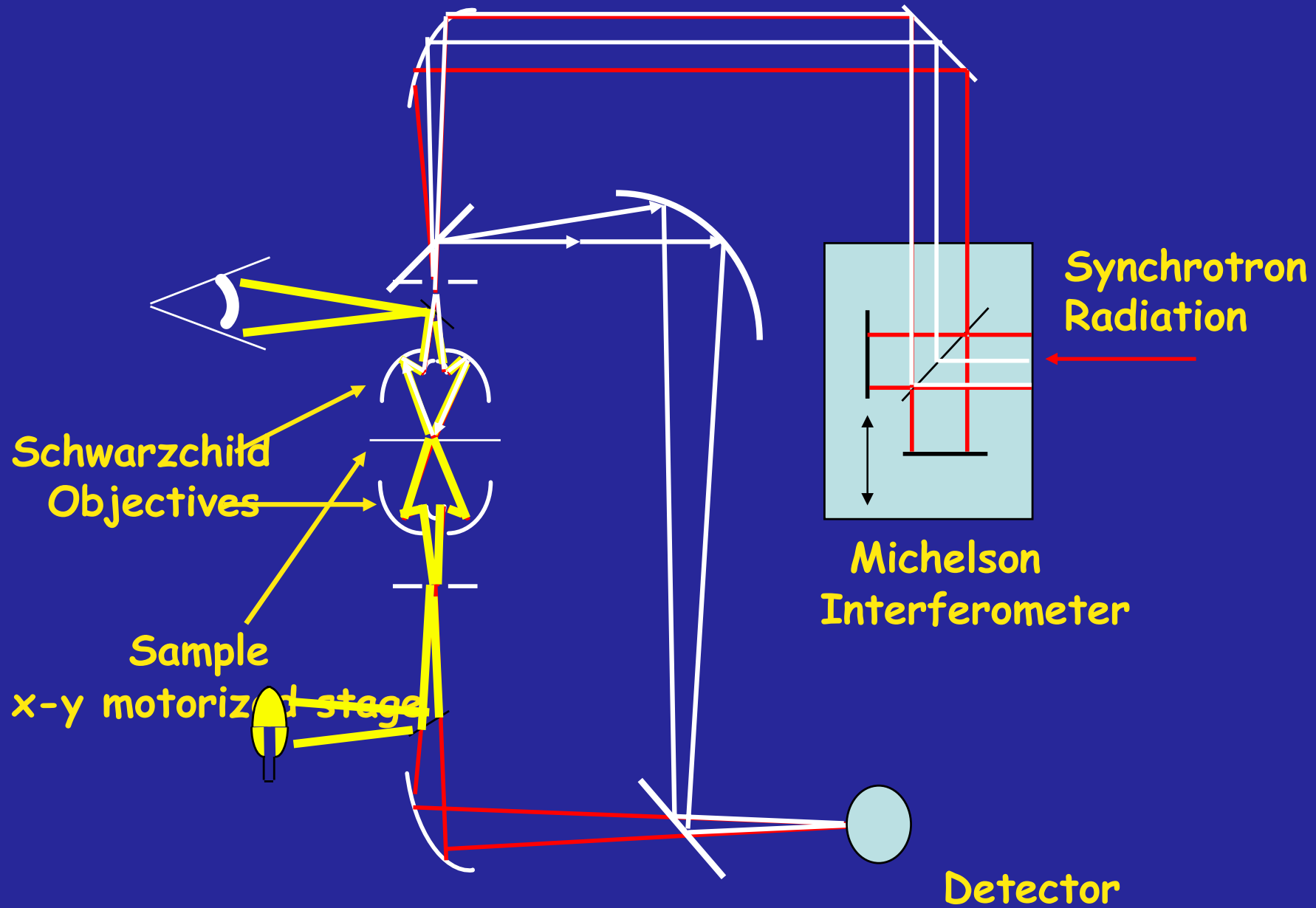
Collimated  
beam  
from  
Storage  
Ring



FTIR  
Spectrometer

Infrared  
Microscope

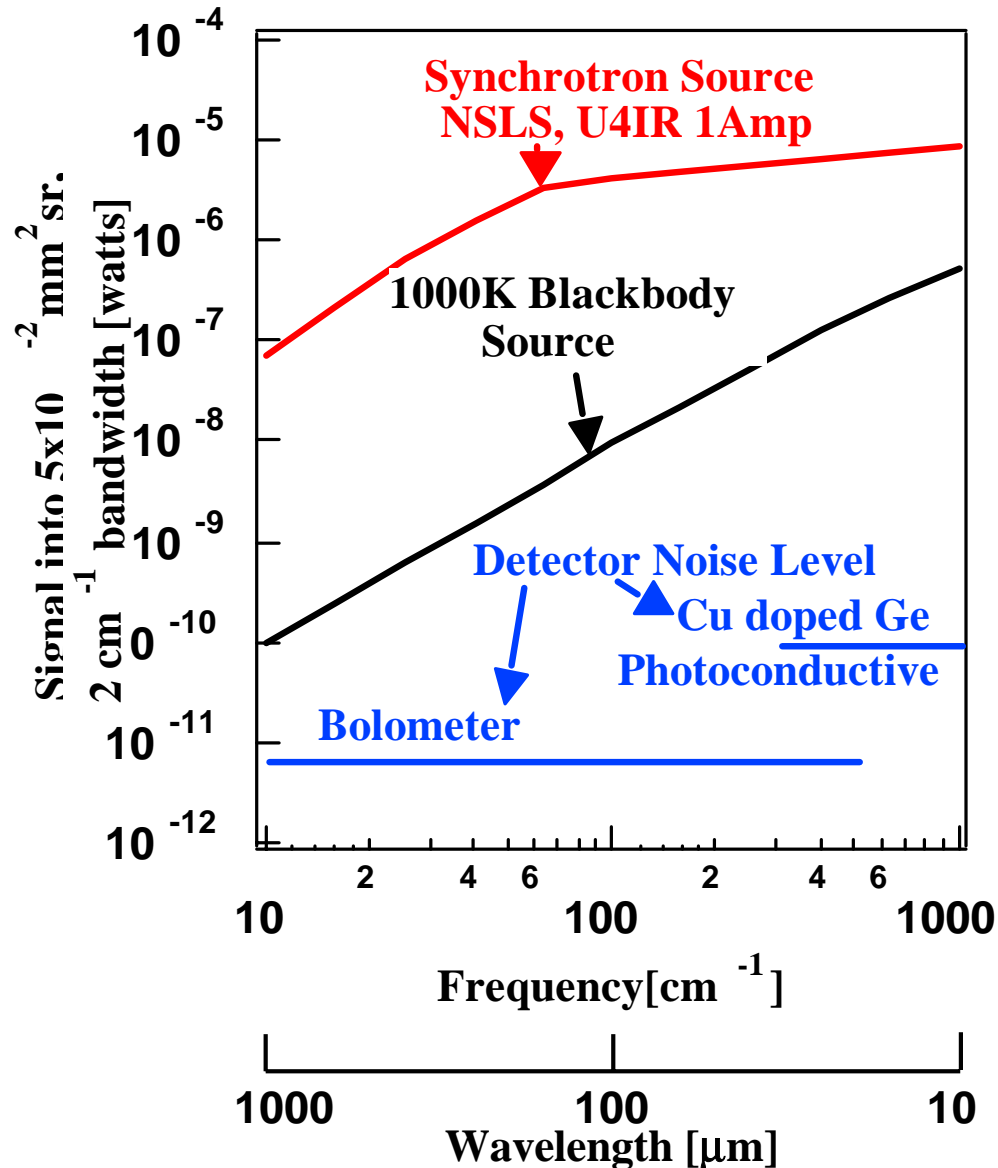
# Infrared Microscope



Infrared  
Beamlines  
U12,  
U10A and  
U10B  
at the  
NSLS



# How Much Signal?



Signal to Noise Ratio:

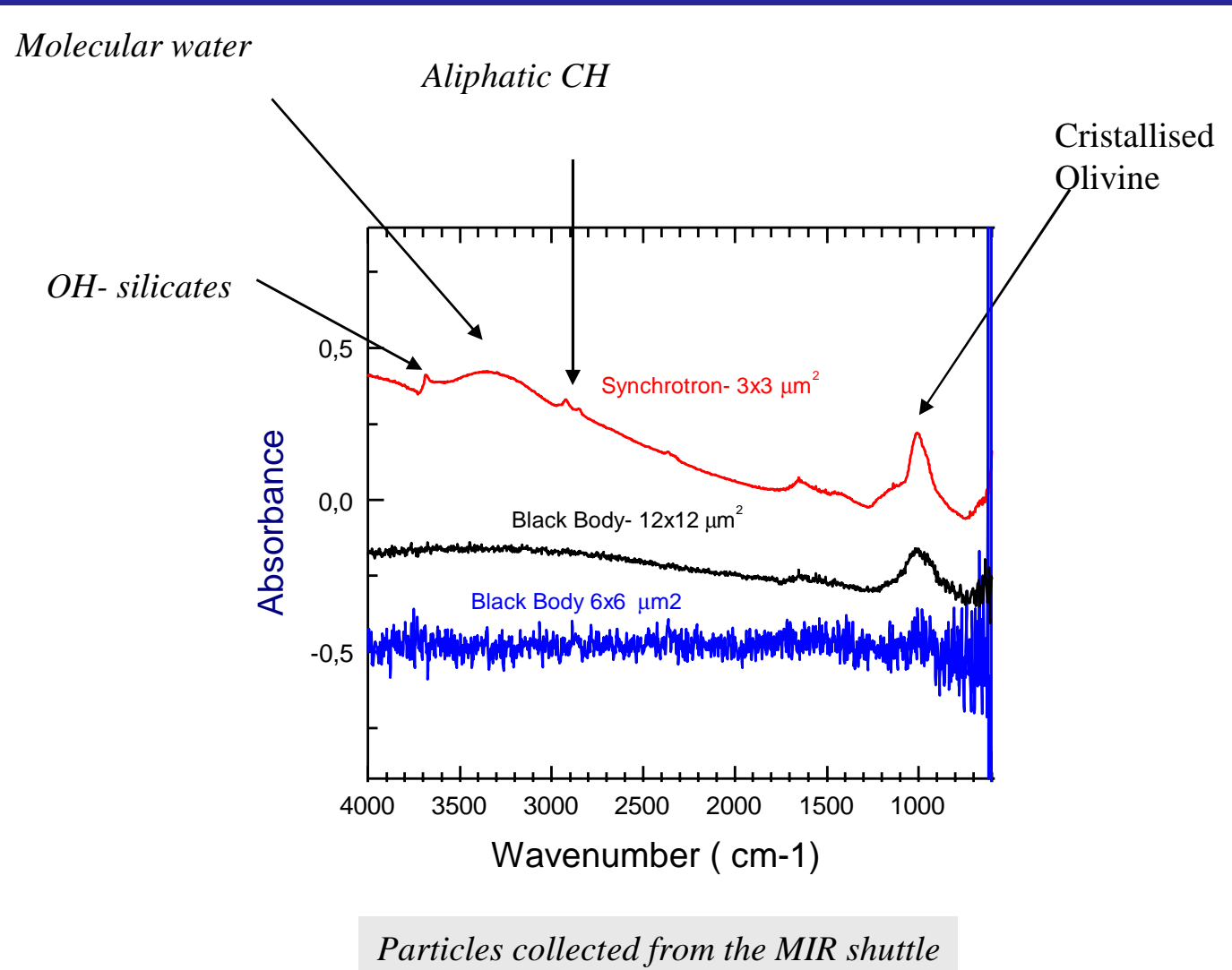
is 100-1000 times larger from a synchrotron source than from a global through a limited aperture

Liquid Helium Cooled detectors to reduce background noise

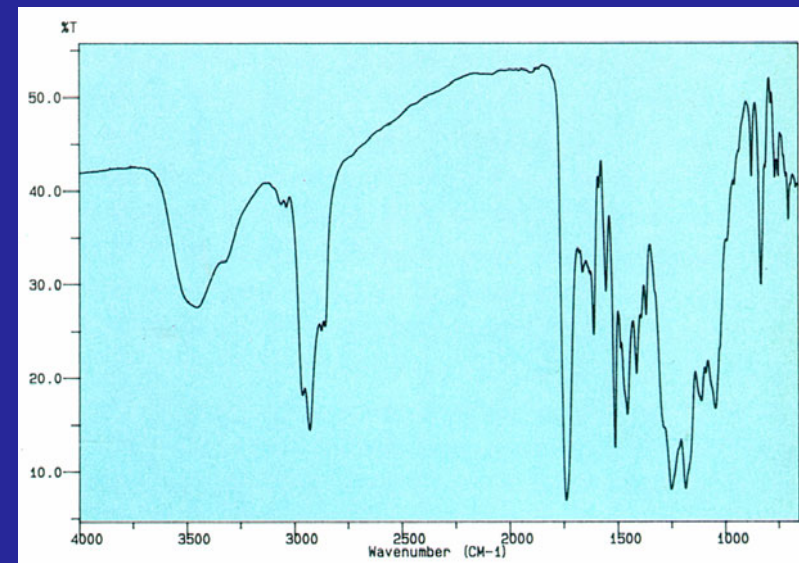
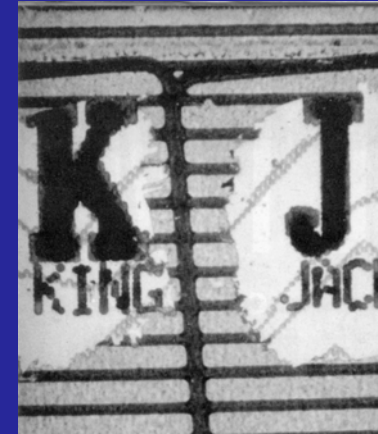
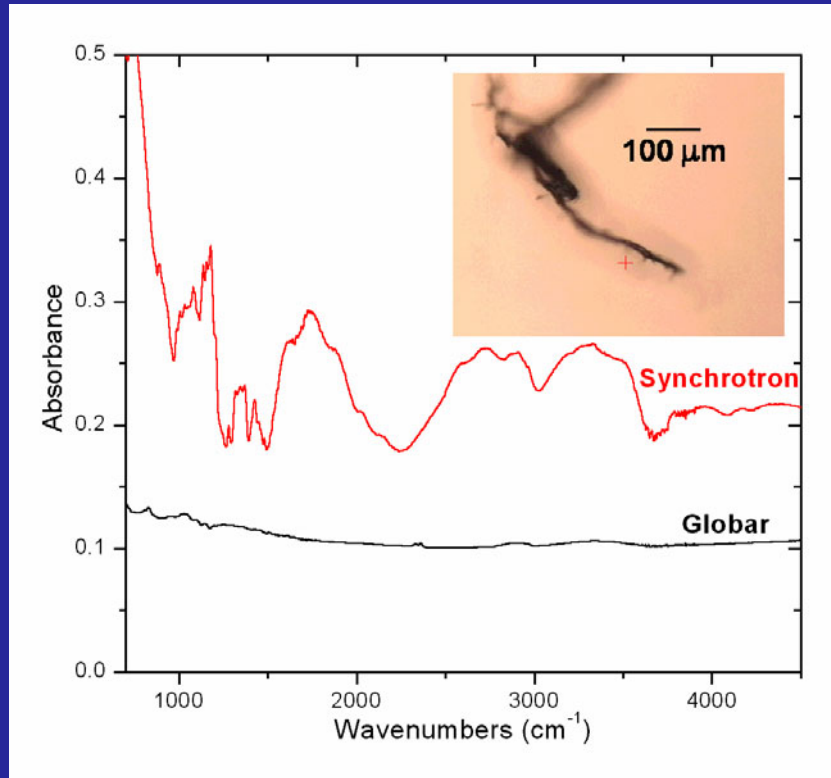
# Interstellar micron size dust analysis

*comparison SRS-Black body:  
3 micron Particles from asteroid "Orgueil"*

From:  
G.Quitte  
J.P.Bibring,  
L.D'Hendecourt  
( IAS),  
P.Dumas ( LURE),  
G.L.Carr,  
G.P. Williams  
( NSLS)



# Chemical Fingerprint of Ink



Wilkinson et al., Physics World, March 2002, 43



# Infrared Microscopy: IR Imaging of Living Cells

## *Euglena gracilis*

Carol Hirschmugl, Maria Bunta, Justin Holt  
*Physics Department, University of Wisconsin-Milwaukee*

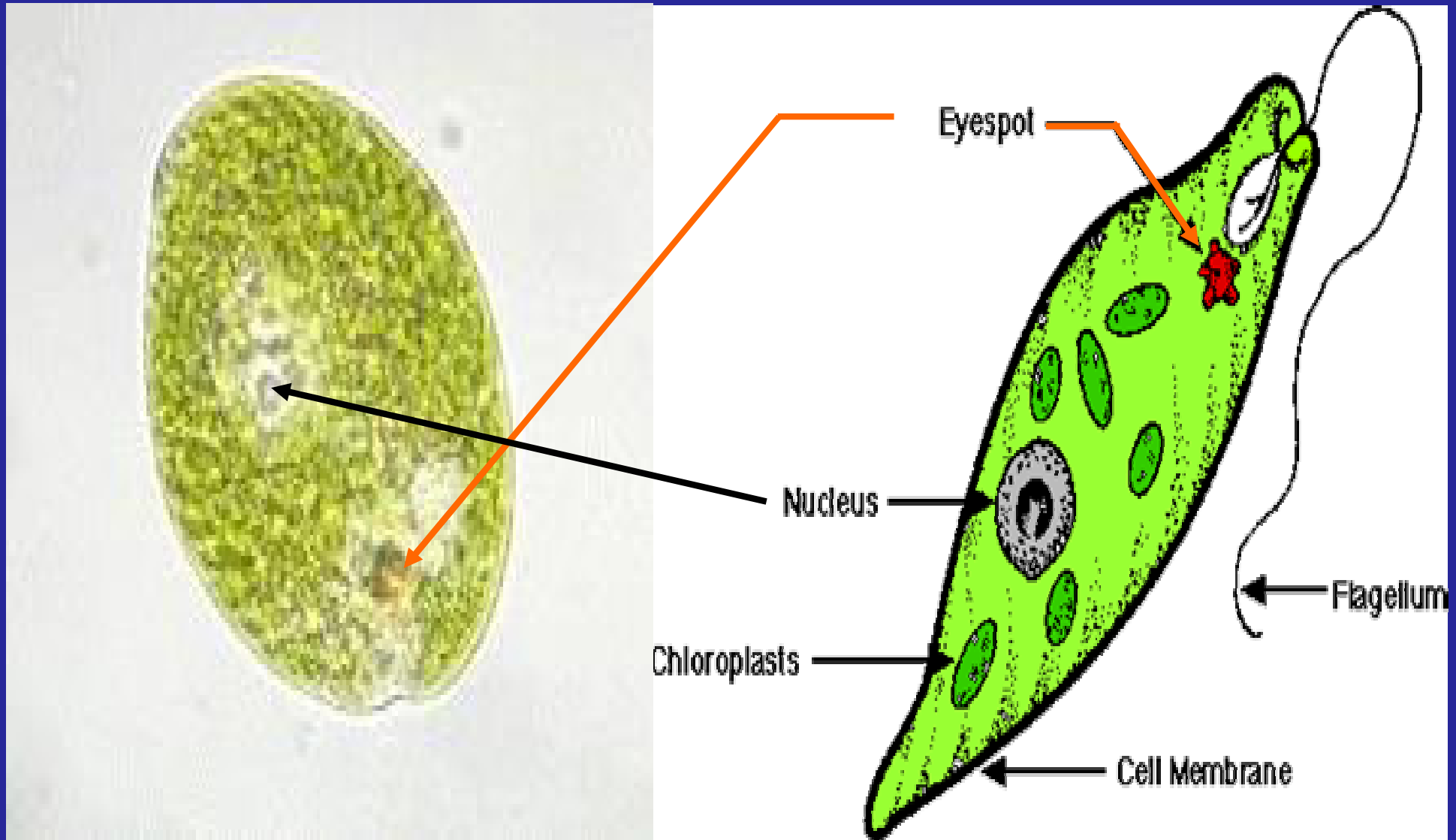
Andrej Skylarov  
*Advanced Analysis Facility, University of Wisconsin Milwaukee*

J Rudi Strickler  
*WATER INSTITUTE, University of Wisconsin Milwaukee*

and  
Mario Giordano  
*Marine Biology, University of Ancona, ITALY*

# Algae Biology : *Euglena gracilis*

approx 15 x 15 microns



# *Euglena gracilis* under the microscope

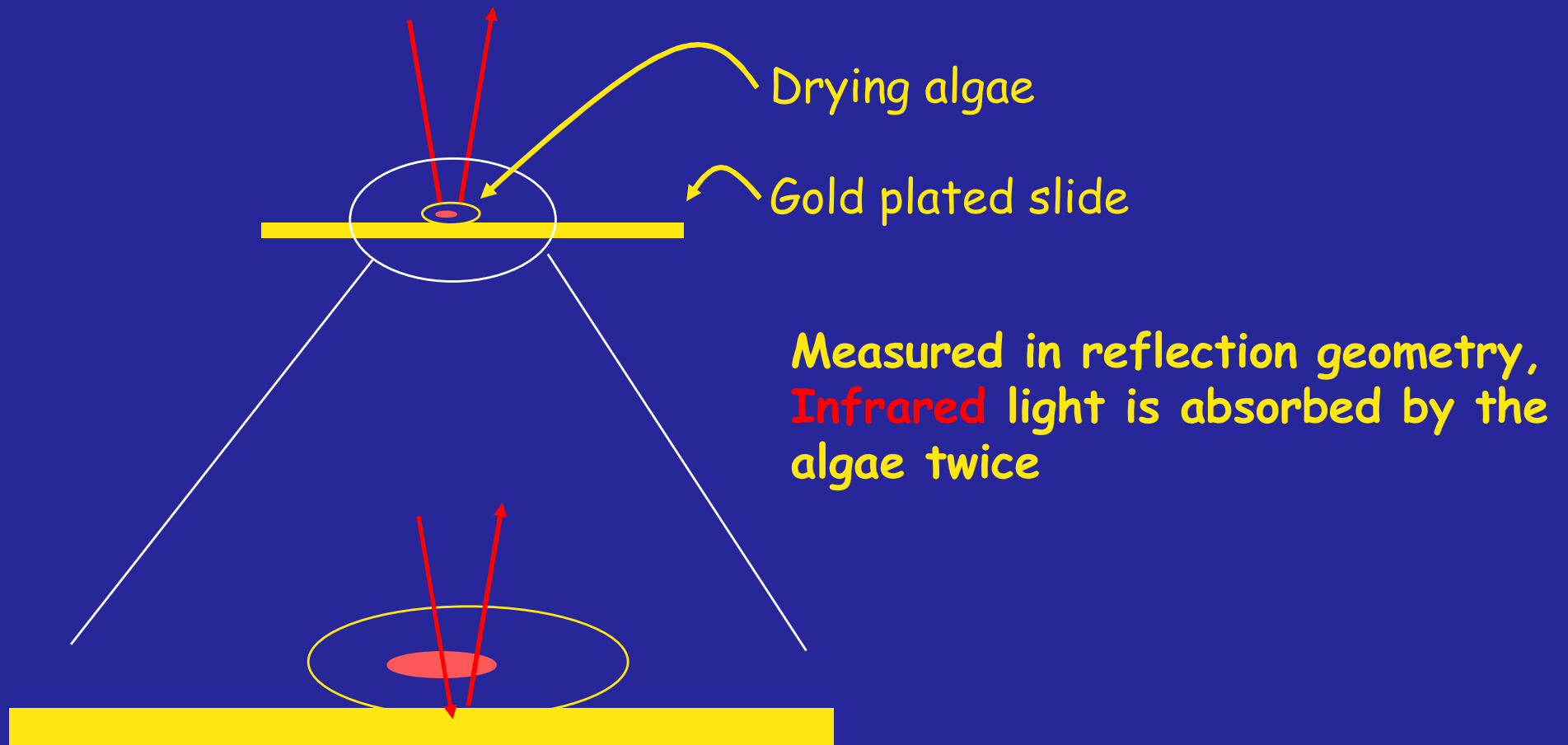


*Euglena gracilis*  
swimming

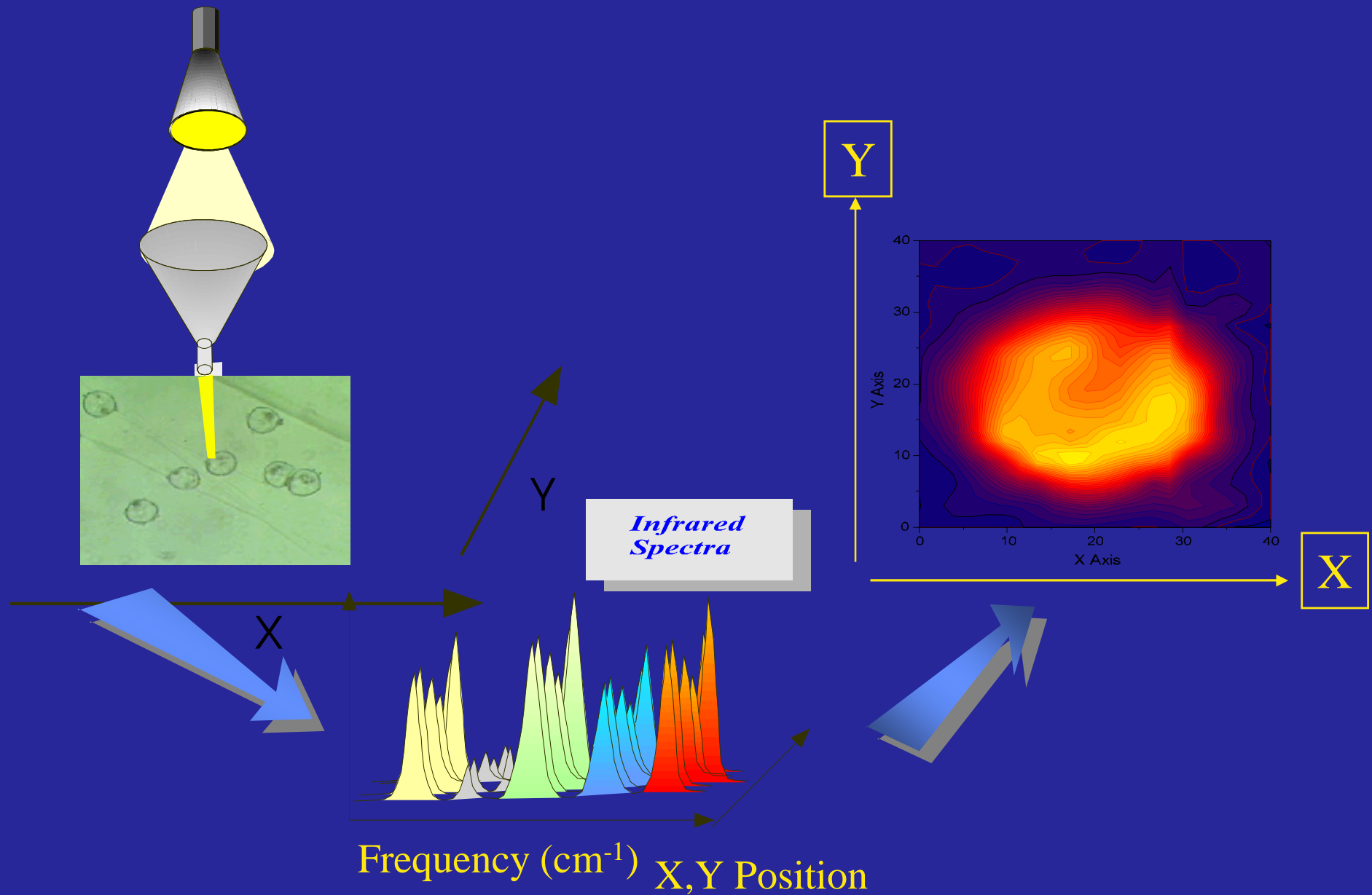


*Euglena gracilis*  
encysted

# Experimental Conditions



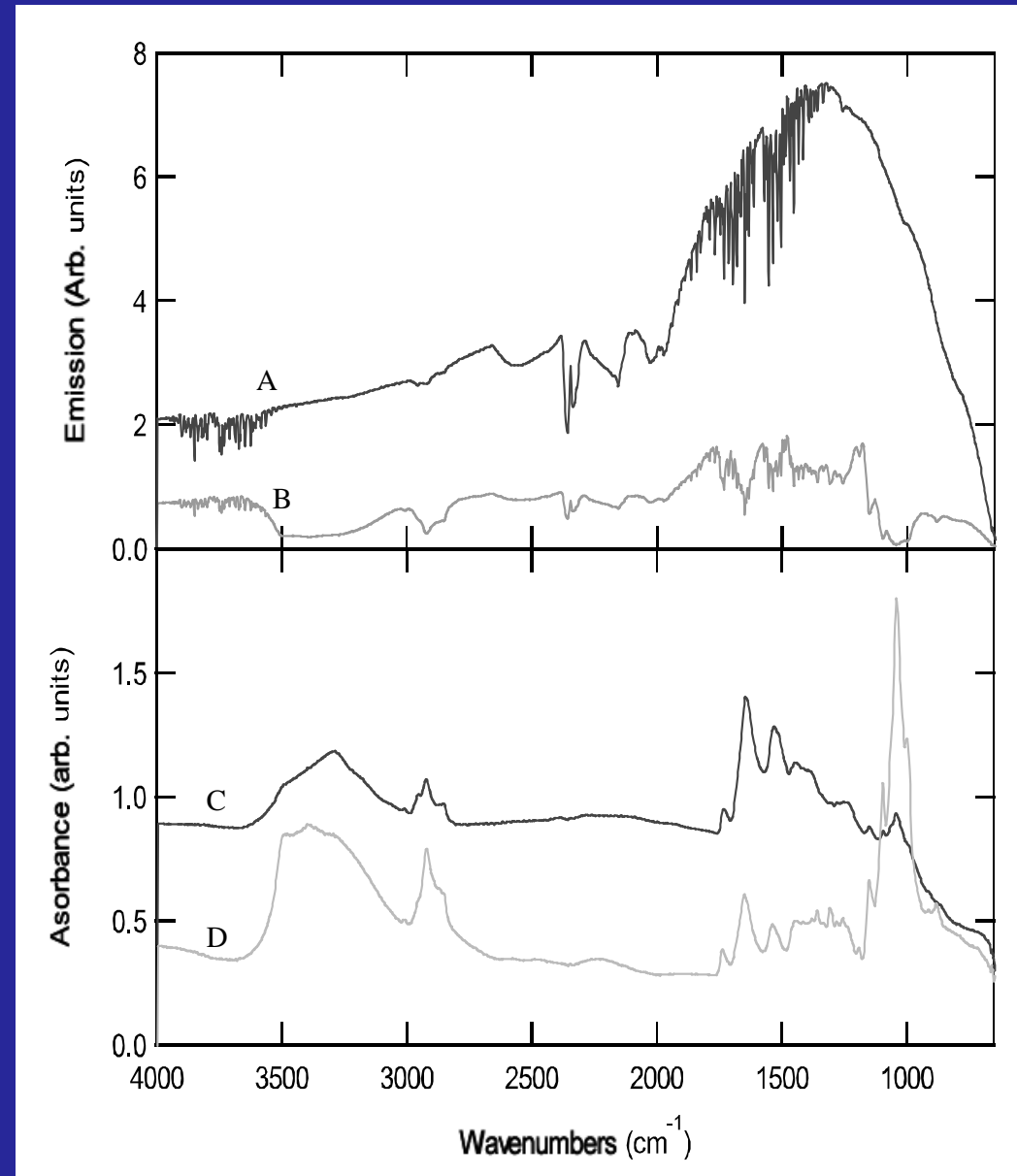
# Chemical Mapping and Chemical Image:



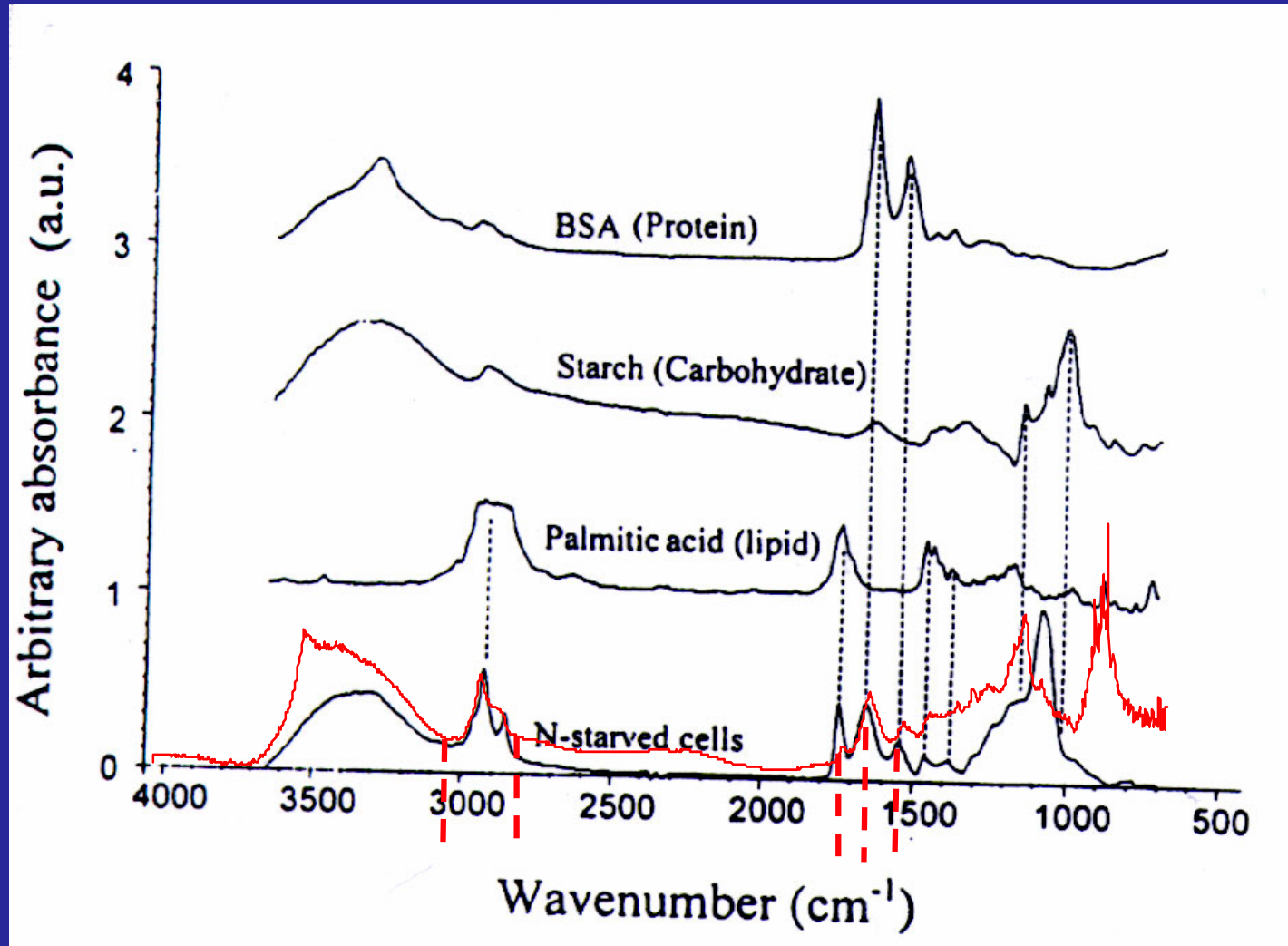
# Reflected and Absorbed Intensities From *Euglena gracilis*

Reflectivity from  
A) Gold  
B) Gold + Alga

Absorbance from  
C) Alga #1  
D) Alga #2



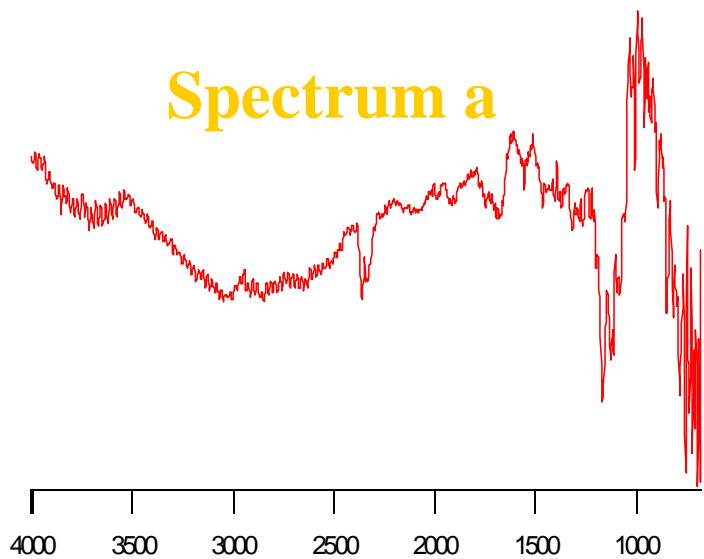
# Infrared Spectrum from Alga and Chemical Standards



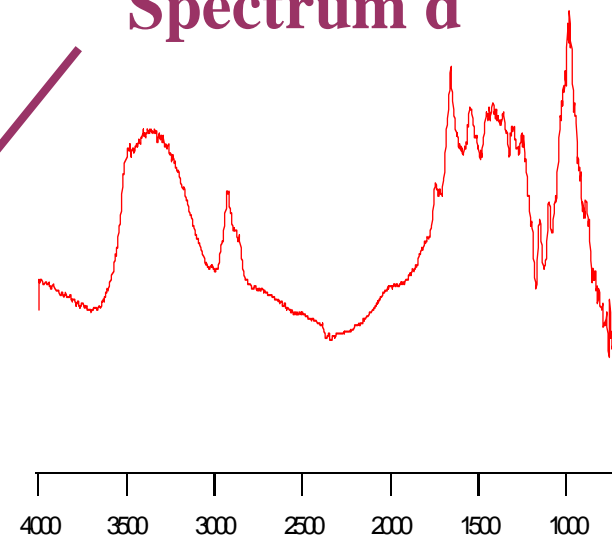
Giordano, et al., *J. of Phycology*, **37**, 271 (2001)

# *Euglena gracilis* and IR spectra

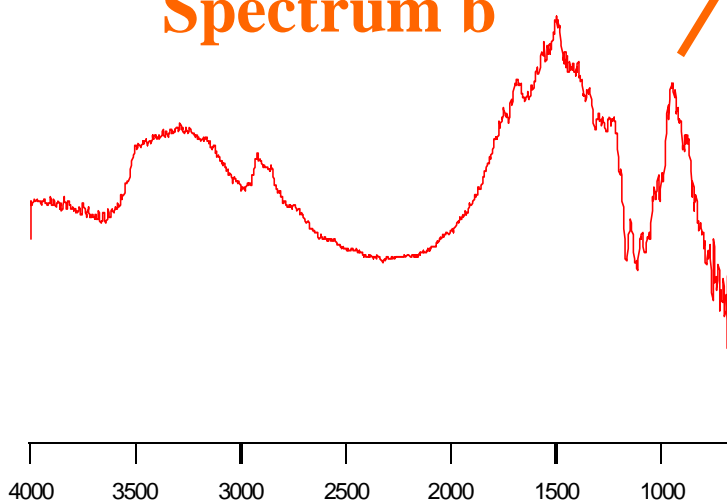
**Spectrum a**



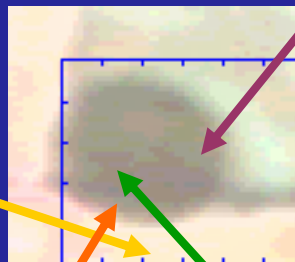
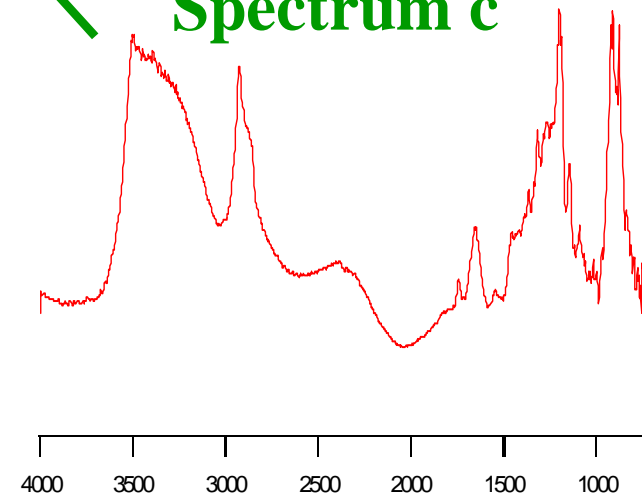
**Spectrum d**



**Spectrum b**

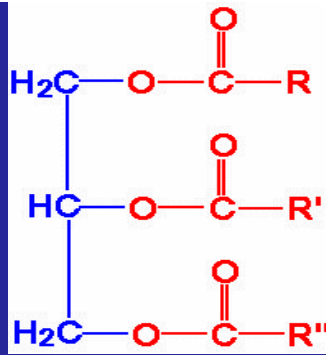


**Spectrum c**





# Lipids



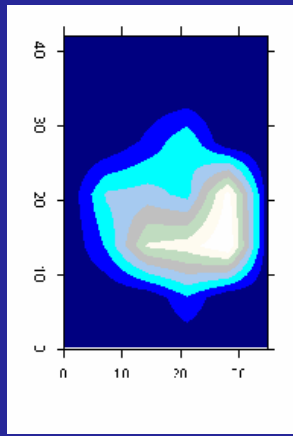
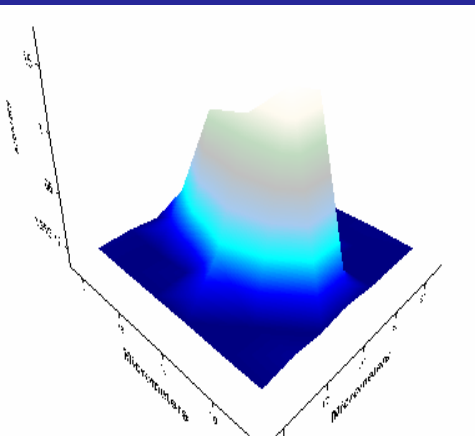
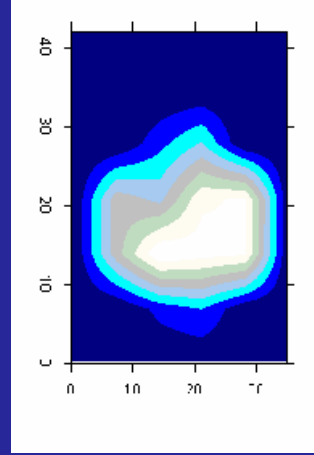
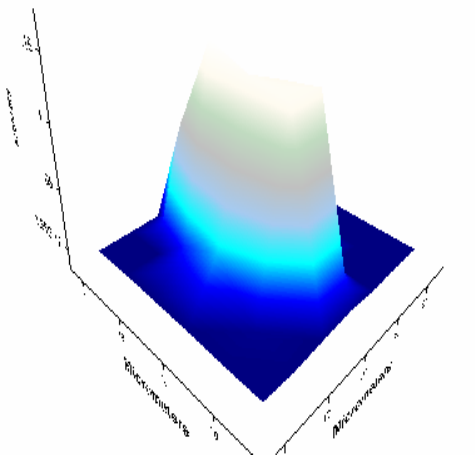
a. C=O functional group



b. CH<sub>2</sub> functional group

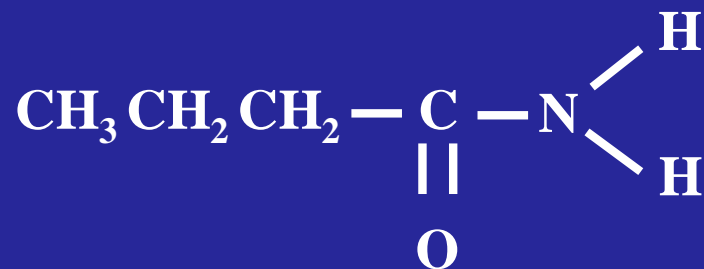


symmetric CH<sub>2</sub> stretch  
vibration at  $2852 \text{ cm}^{-1}$

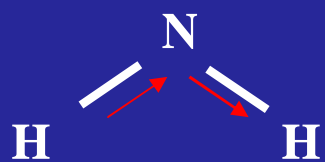


# Proteins

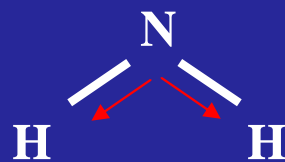
## Amides I



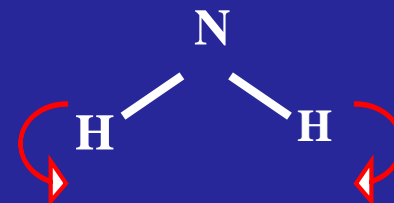
### a. $\text{NH}_2$ functional group



Asymmetric  
Stretch

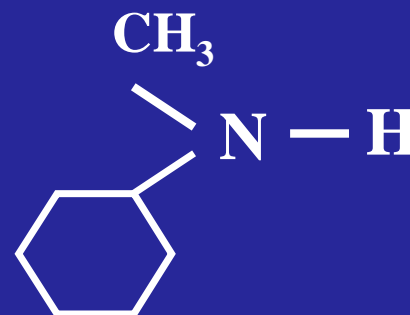


Symmetric  
Stretch



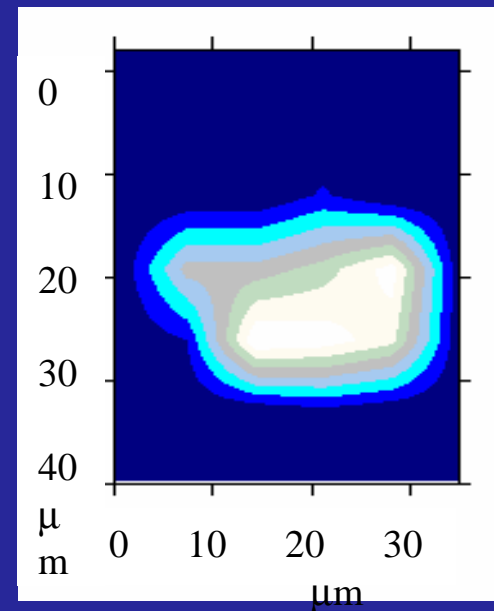
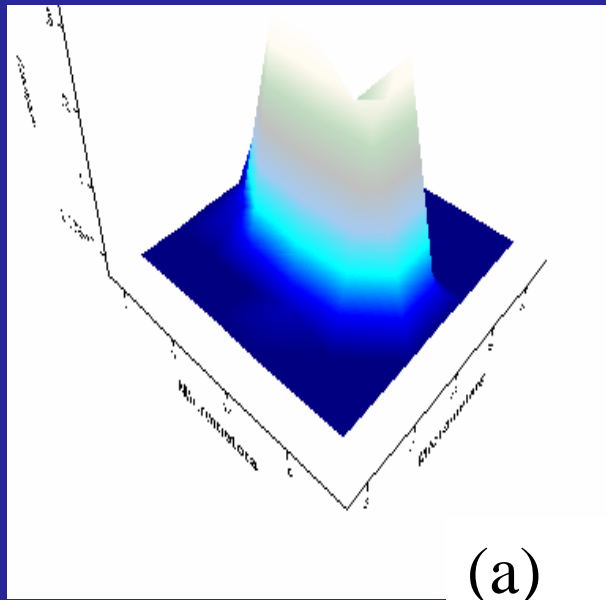
Scissors  
Mode

## Amides II

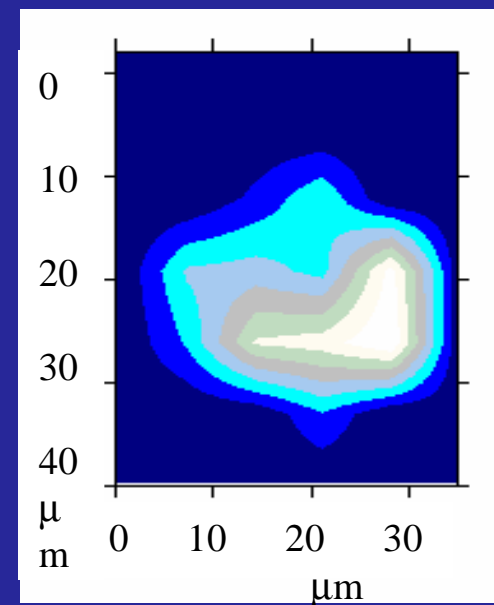
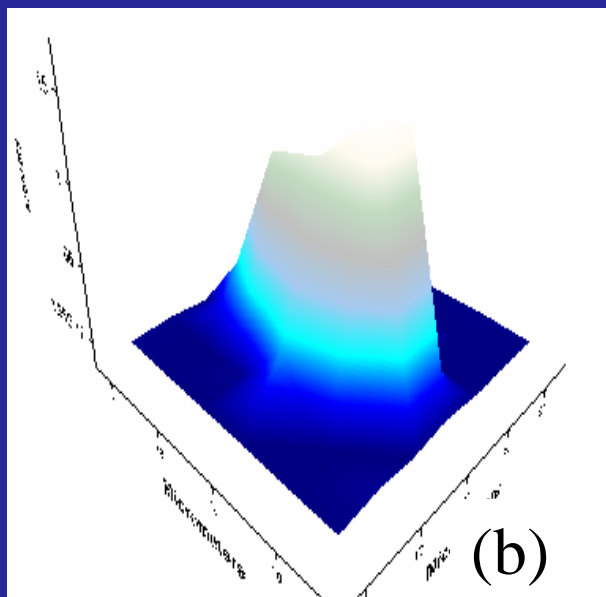


- Amides I peak at  $1650\text{ cm}^{-1}$  represents 80%  $\text{C}=\text{O}$  stretching vibration, 10%  $\text{C}-\text{N}$  stretching vibration and 10%  $\text{N}-\text{H}$  bending
- Amides II peak at  $1545\text{ cm}^{-1}$  represents 60%  $\text{N}-\text{H}$  bending and 40%  $\text{C}-\text{N}$  stretching vibration

# *Euglena gracilis* and Chemical Map: Proteins



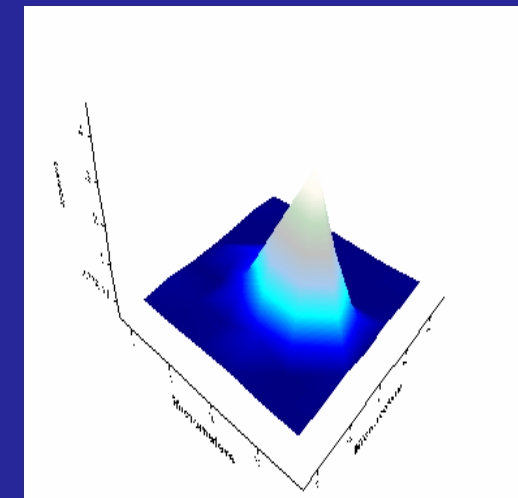
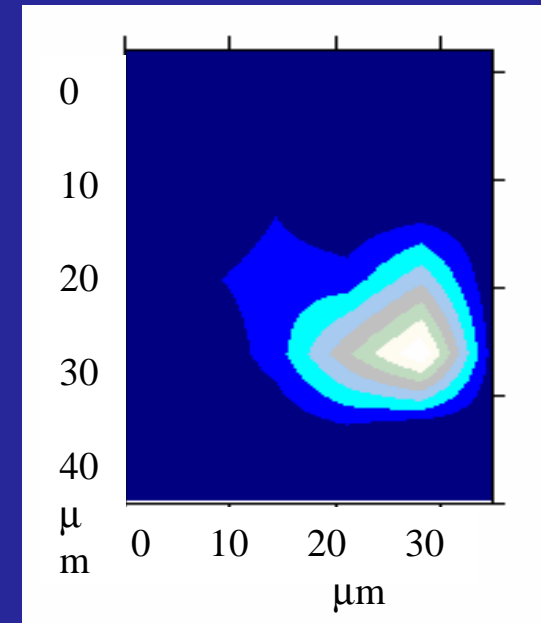
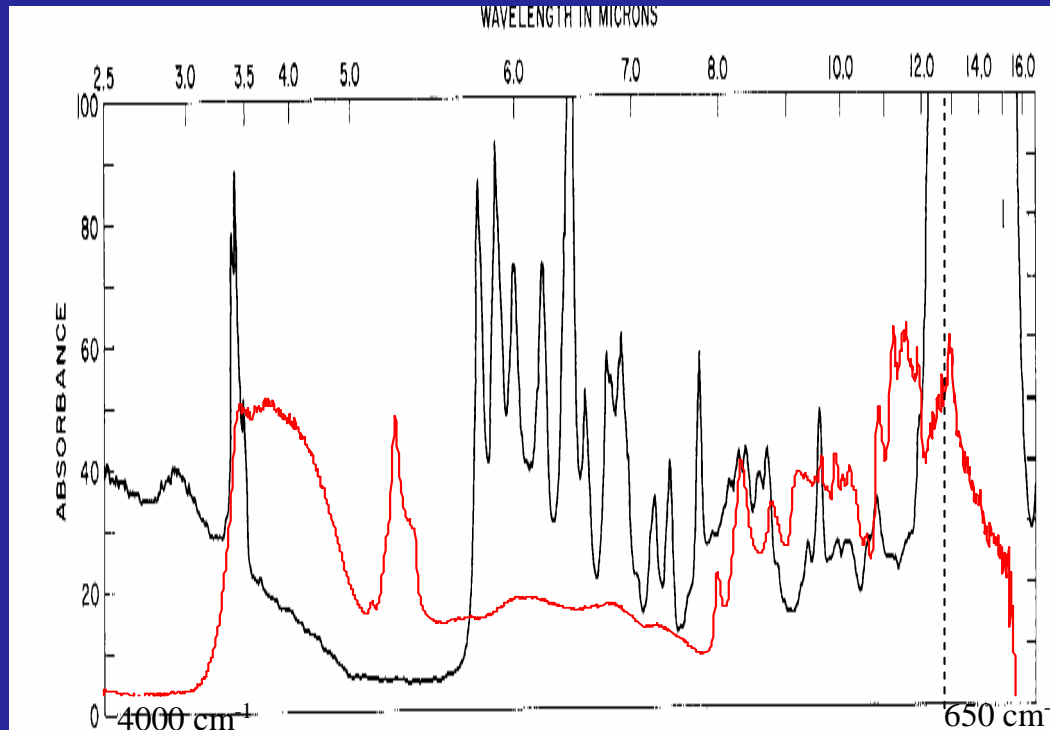
AMIDES I +  
Water  
Peak 1650cm<sup>-1</sup>



AMIDES II  
Peak 1545 cm<sup>-1</sup>



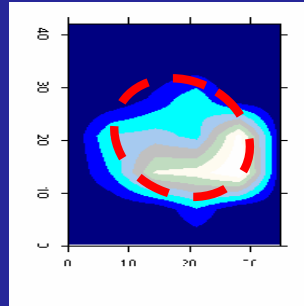
# Chlorophyll Spectrum



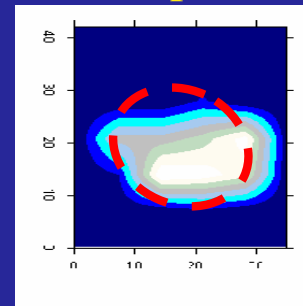
Vernon LP, Seely GR, *The Chlorophylls*, Academic Press, New York, London, 1966, 193.

# Summary of Findings

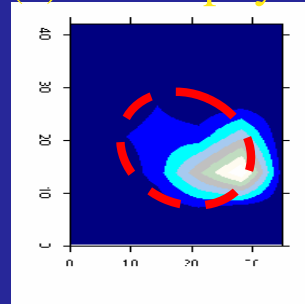
(a) Protein



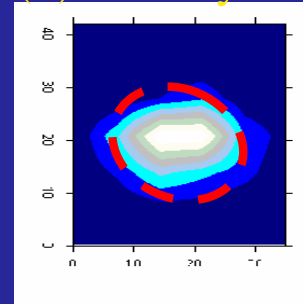
(b) Lipid



(c) Chlorophyll

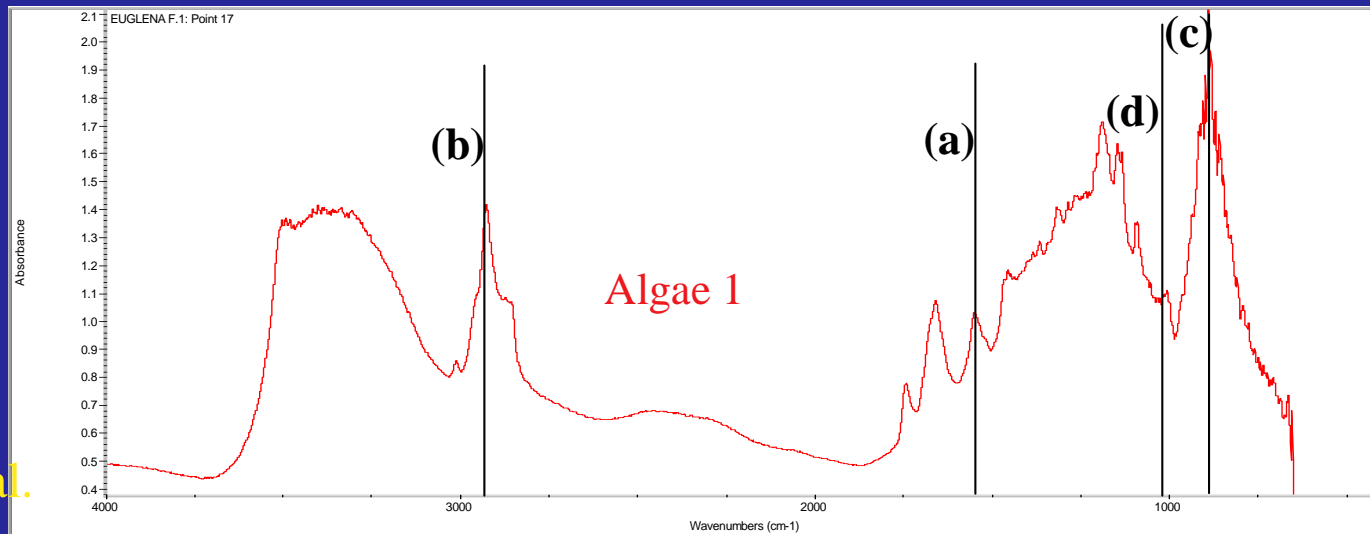


(d) Paramylon

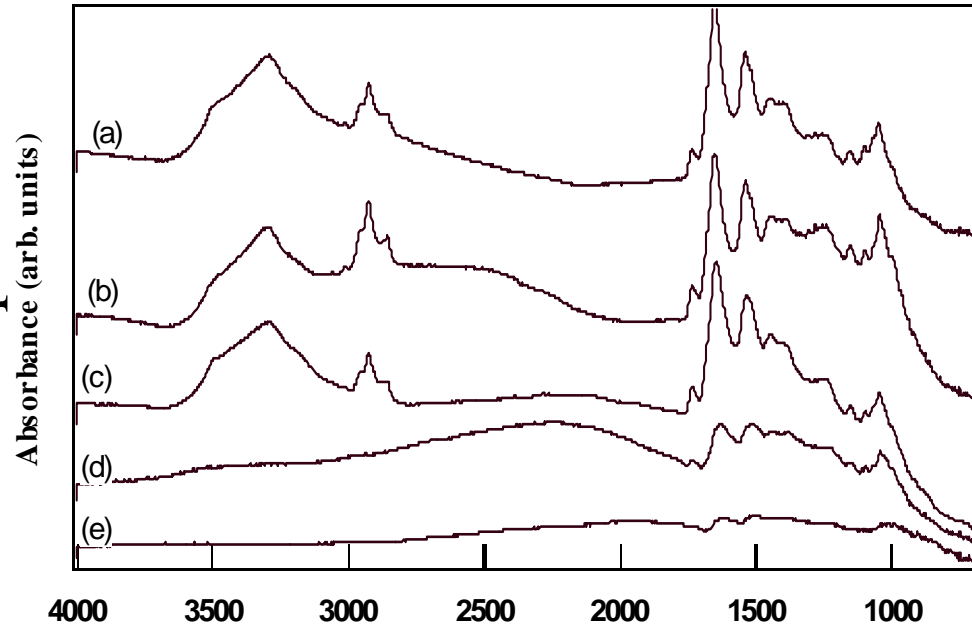


6  $\mu\text{m}$  x 6  $\mu\text{m}$   
spectrum at  
chlorophyll  
maximum

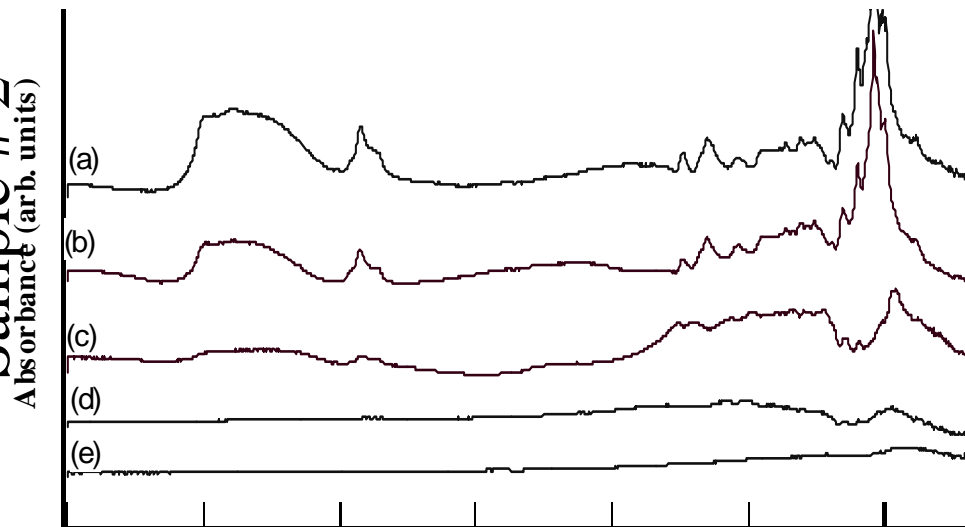
Hirschmugl et. al.



## Sample # 1



## Sample # 2

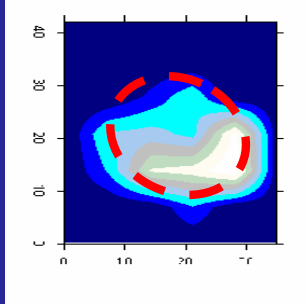


- the protein to carbohydrates ratio :
  - sample # 1 alga is 1:0.2
  - sample # 2 alga is 1:10
- the relative strength of the lipid to protein band dramatically different
- Sample #2 is nutrient starved
- Conclude: sample # 2 is nitrogen starved

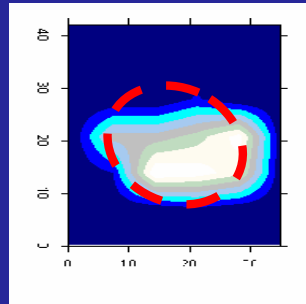
# Summary for 2 Samples: *Euglena Gracilis*

## Alga #1

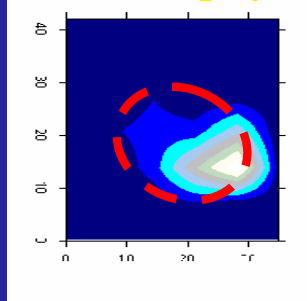
(a) Protein



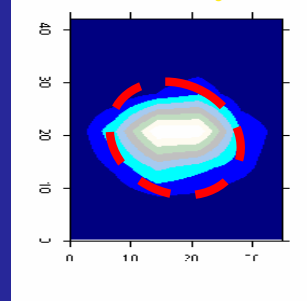
(b) Lipid



(c) Chlorophyll

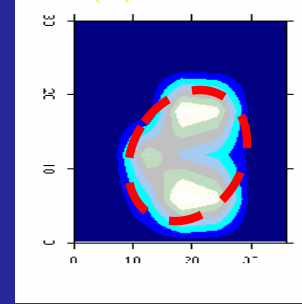


(d) Paramylon

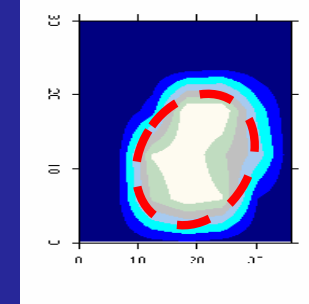


## Alga #2

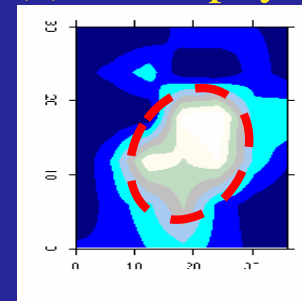
(a) Protein



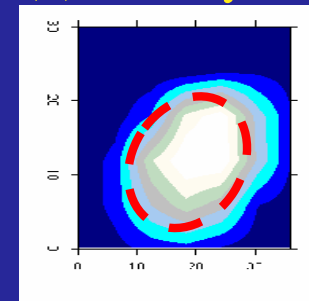
(b) Lipid



(c) Chlorophyll

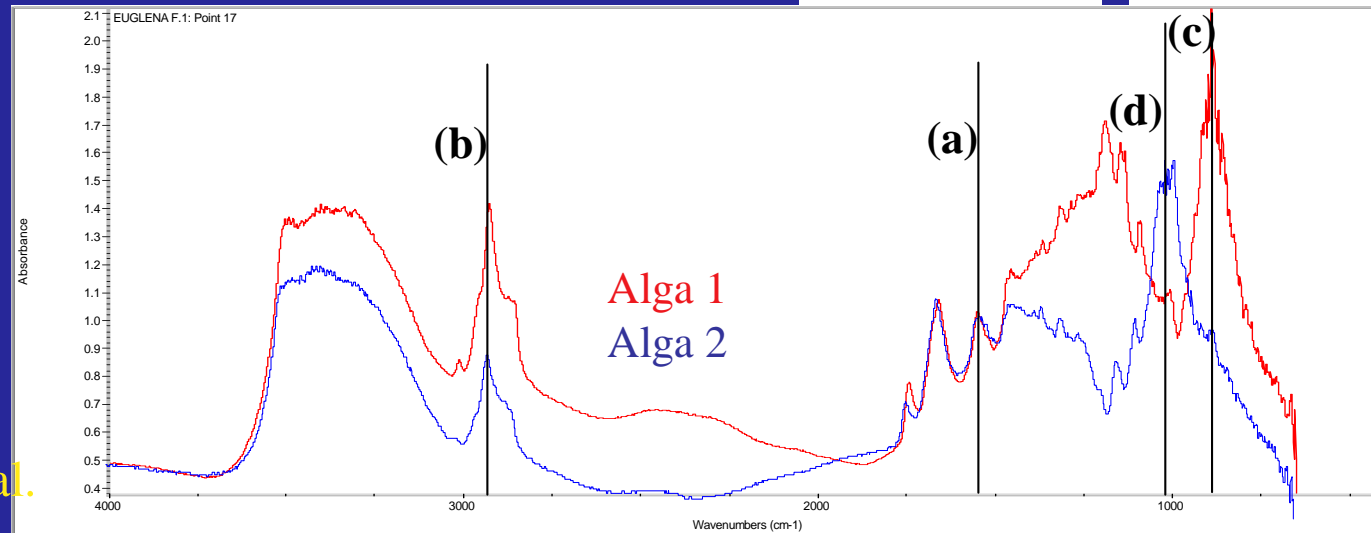


(d) Paramylon



6  $\mu\text{m}$  x 6  $\mu\text{m}$   
spectra at  
chlorophyll  
maxima

Hirschmugl et. al.



# Summary: Algae

IR absorption contour maps of individual alga obtained

Contour maps of individual spectral features show similar structure as single cell

Chemical Features for two separate alga agree with Nitrogen Starvation Model

## Future

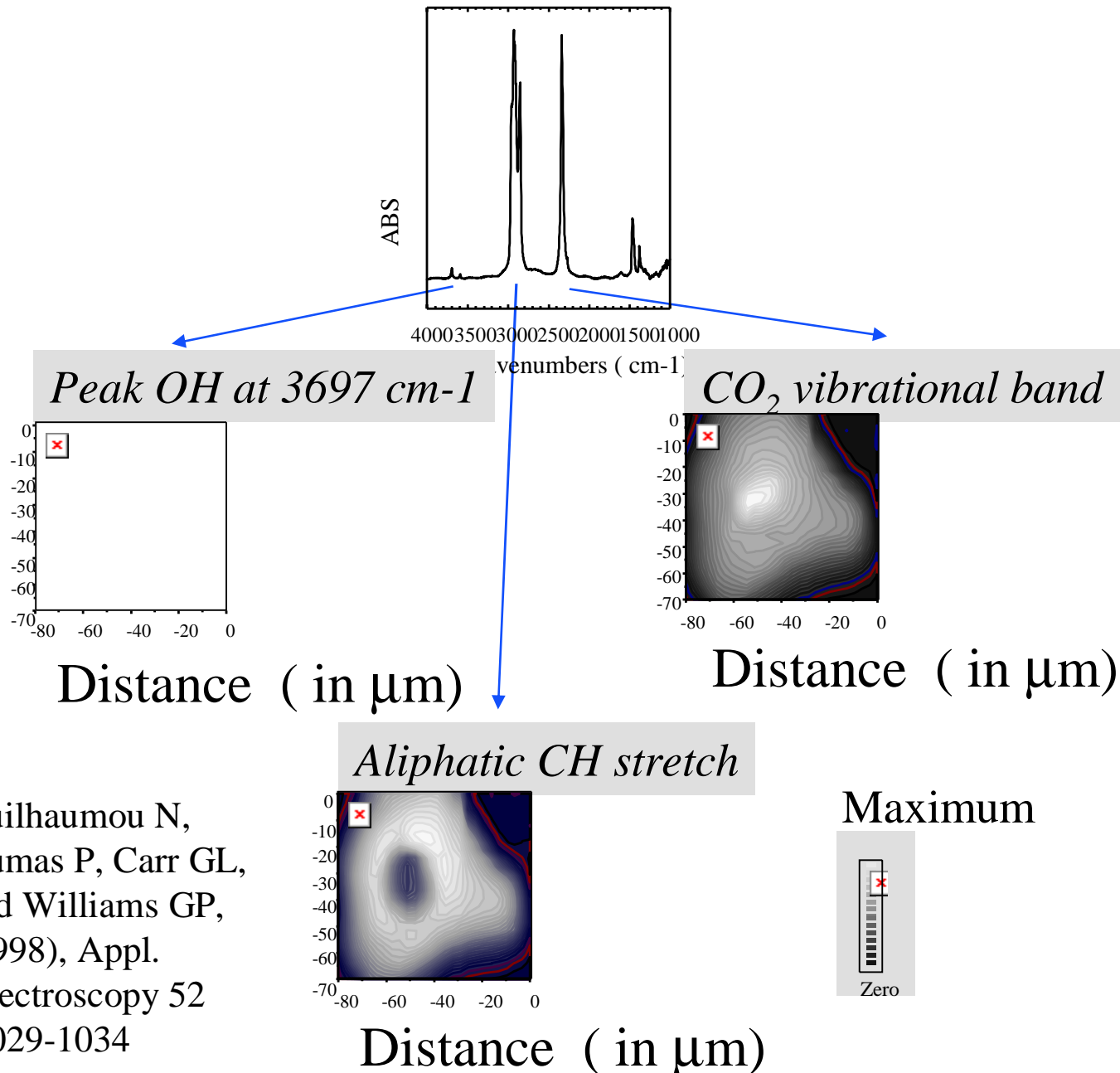
Identify the origin of other ir features

Measure different algae species under controlled nitrogen environments



# Chemical Nature: Inclusions in Geological Specimen

(Guilhaumou et al.)



Guilhaumou N,  
Dumas P, Carr GL,  
and Williams GP,  
(1998), Appl.  
Spectroscopy 52  
:1029-1034

**High Pressure  
Gas Phase  
CO<sub>2</sub>**

**Aliphatic  
Hydrocarbon  
(Chains)  
Liquid Phase  
Terminated with  
Hydroxyl Groups**

# Interplanetary Dust Particles (IDPS)

(Bradley et al.)

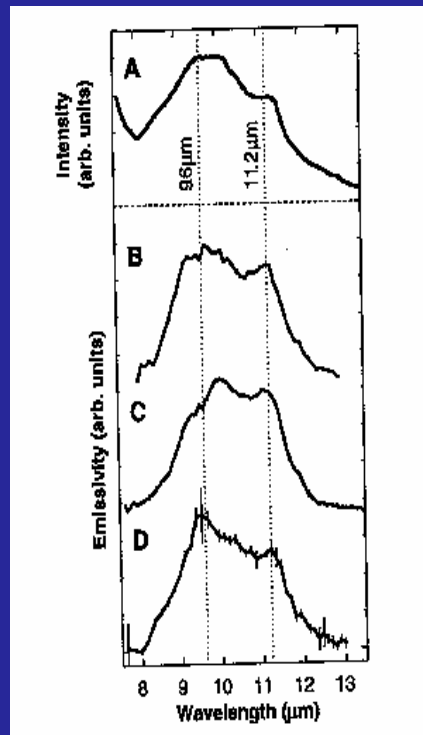
## Si-O Stretch Region at 10 $\mu\text{m}$

Enstatite 9.6  $\mu\text{m}$  + Fosterite 11.2  $\mu\text{m}$   
+ glassy silicates (broad)

Amorphous silicates

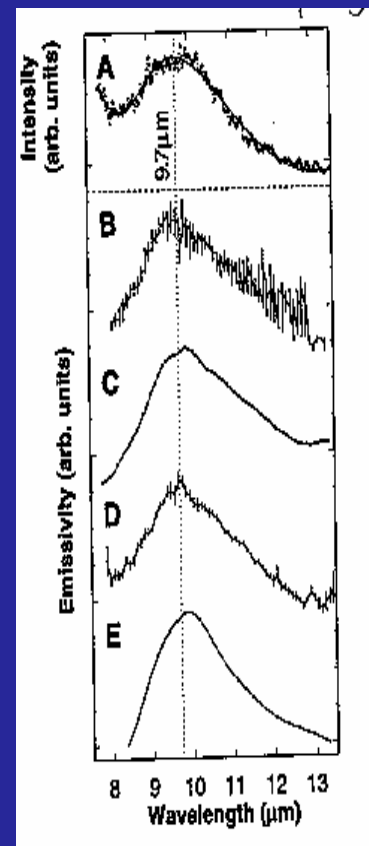
Absorption Spectrum  
A: GEMS

Emission Spectra  
B: Elias Molecular Cloud  
C: Trapezium Molecular Cloud  
D: T Tauri Young stellar object DI Cephei  
E: M type super gienat  $\mu$ -Cephei



Absorption Spectrum  
A: GEMS rich IDP

Emission Spectra  
B: Comet Halley  
C: Comet Hale-Bopp  
D: Late Stage Herbig star



# Interplanetary Dust Particles (IDPS)

(Bradley et al.)

## Conclusions:

Presolar Interstellar Molecular Cloud could consist of GEMS

Presume solar system formed from intermolecular cloud.

Long sought building block of Solar System.

# Environmental Remediation of Soil

(Ghosh et al.)

## Goal:

Examine Aliphatic and Polycyclic Aromatic Hydrocarbons (PAH) sorbed onto different components in sediment  
e.g. silica, coal wood

## Experimental methods (combination of techniques):

IR microspectroscopy

Scanning Electron Microscopy

Laser desorption/laser ionization mass spectrometry

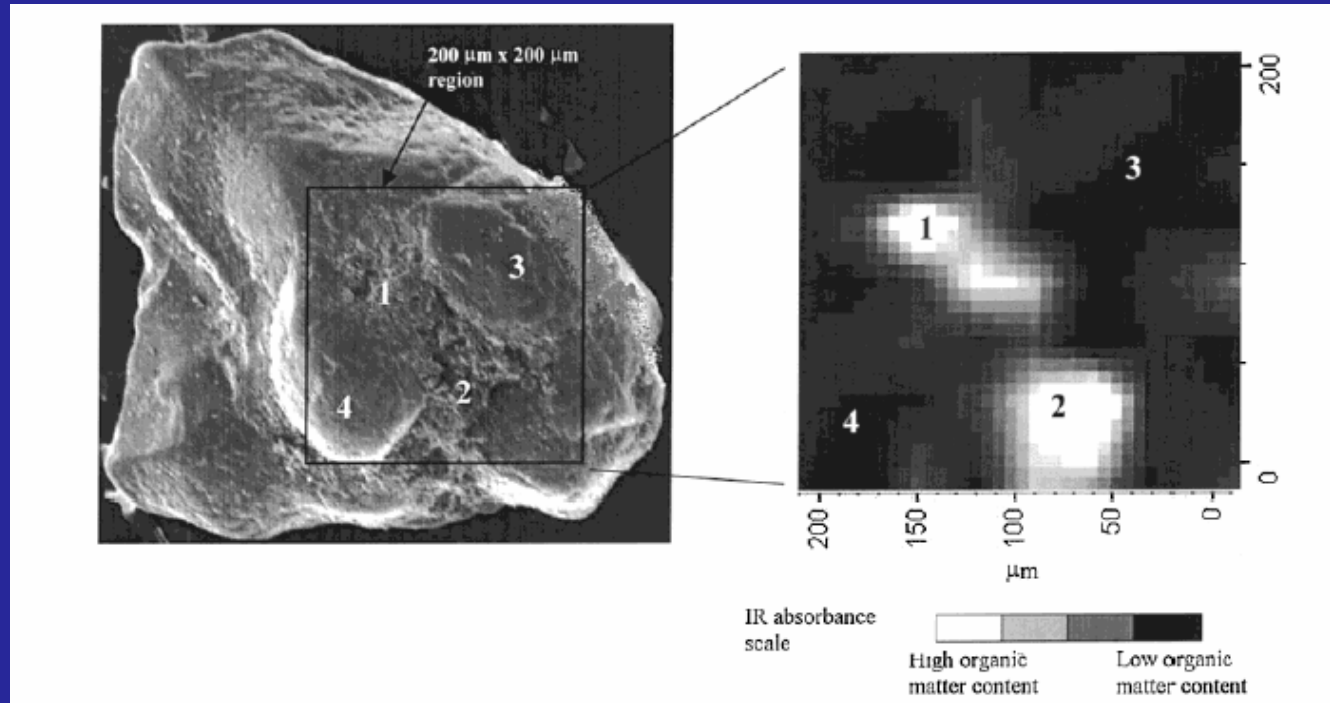
Aliphatic HCs: linear chains of hydrocarbons

C-H stretches  $\Rightarrow$  2800-3000  $\text{cm}^{-1}$

Aromatic HCs: cyclic hydrocarbons

C-H stretched  $\Rightarrow$  above 3000  $\text{cm}^{-1}$

# Environmental Remediation of Soil



Scanning Electron Microscope (SEM) image of a silica particle having patches of organic matter as indicated by the white regions in the IR mapping of the C-H stretching absorbance ( $2800 - 3000 \text{ cm}^{-1}$ ) shown in the right panel. (Ghosh et al.)

## Conclusions:

### Silica

Aliphatic HCs are sorbed in patches

Polyaromatic are sorbed in coexistence with Aliphatic HCs

### Coal and Wood

Aliphatic are uniformly distributed throughout

Polyaromatic are concentrated at surfaces, with higher concentration than on silica

# Bone Composition: Phosphate Bands

(Miller et al.)

## Goal:

Determine acid phosphate content and mineral crystallite perfection next to an osteon in bone.

## Experimental methods (combination of techniques):

IR microspectroscopy

X-Ray powder diffraction

Correlation between synthetic hydroxyapatite crystals and natural bone powders of various species and ages.

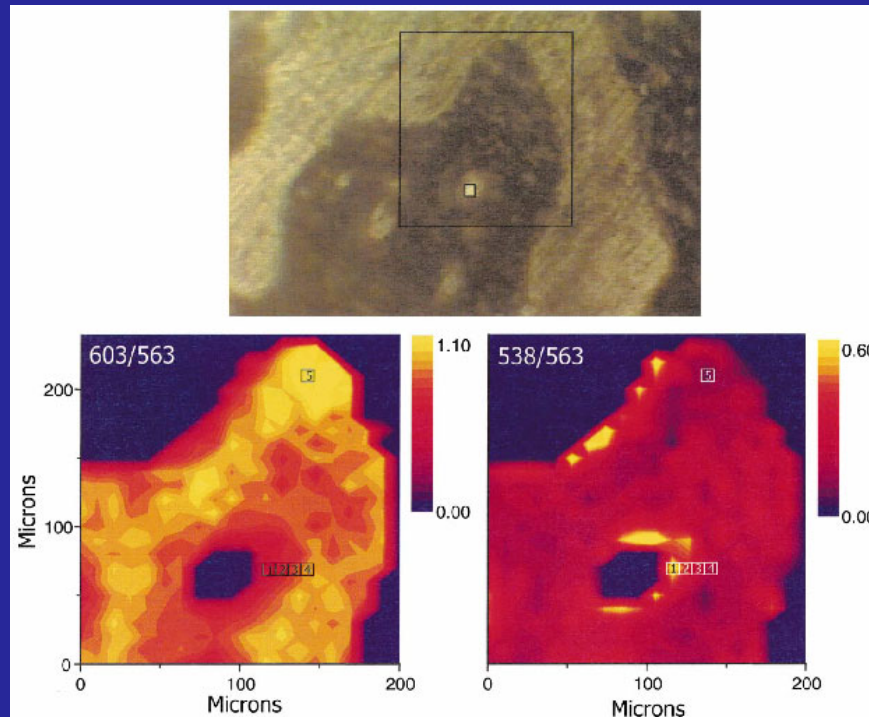
## Phosphate bands ( $\text{PO}_4^{3-}$ ) between 500-650 $\text{cm}^{-1}$ :

603/563 (ratio of absorption strengths)  $\Rightarrow$  Crystallite content

538/563 (ratio of absorption strengths)  $\Rightarrow$  Acid Phosphate content

# Bone Composition: Phosphate Bands

(Miller et al.)



Optical and infrared images of human, osteopetrotic, trabecular bone. Data were collected at Beamline U4IR at the NSLS using a Spectra Tech Ir $\mu$ s microscope and a Cu doped Ge detector (Infrared Laboratories). Apertures were set at 12 by 12  $\mu$ m and 128 scans were collected at 4  $\text{cm}^{-1}$  resolution. Infrared images were generated by plotting peak height ratios of (left) 603/563  $\text{cm}^{-1}$  and (right) 538/563  $\text{cm}^{-1}$ . Miller et al.

## Conclusions:

As bone matures average crystallite size increases  
Acid Phosphate is found in new bone and decreases as bone ages  
Thus, the acid phosphate content and the crystallinity are inversely related.

The most dramatic changes in osteon composition occur within 30 microns of the site of the osteon center.

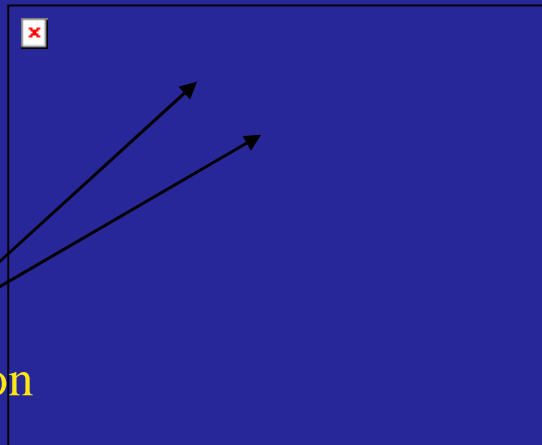
# Chemical Imaging of Living Cells: Dividing Cell

(Jamin et al.)

Optical Image

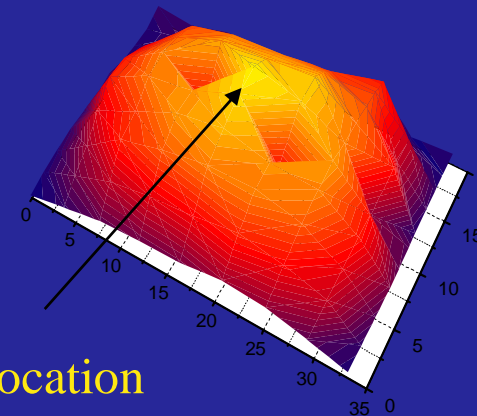


Amide II  
Nucleus



Nucleus location

CH Aliphatic  
Lipids



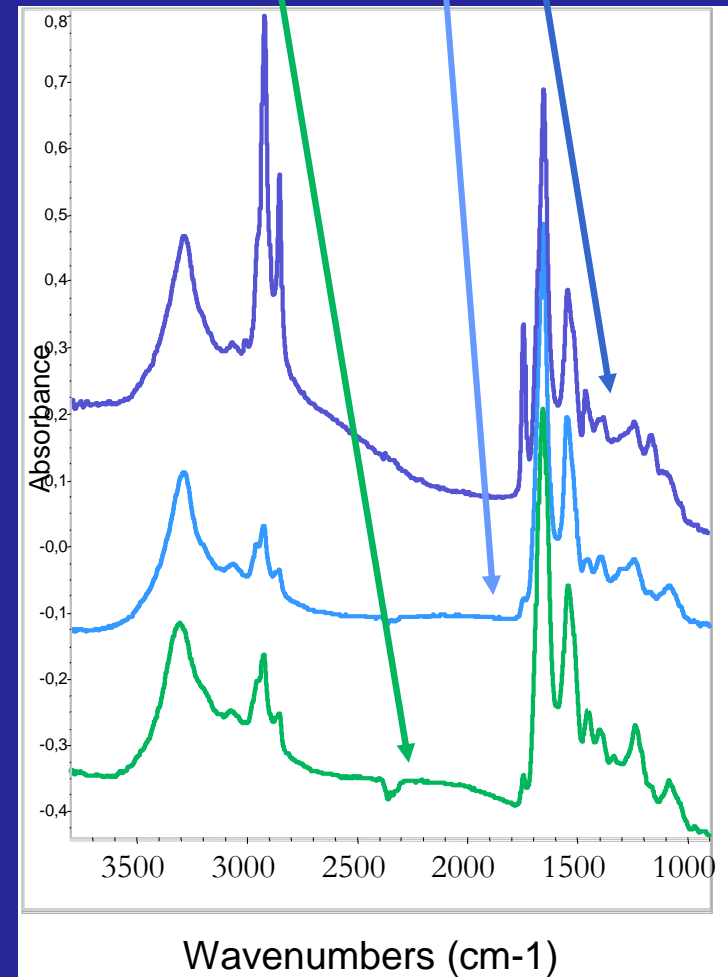
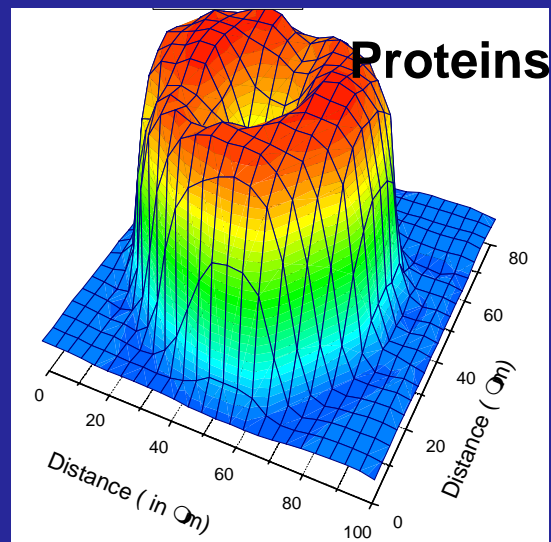
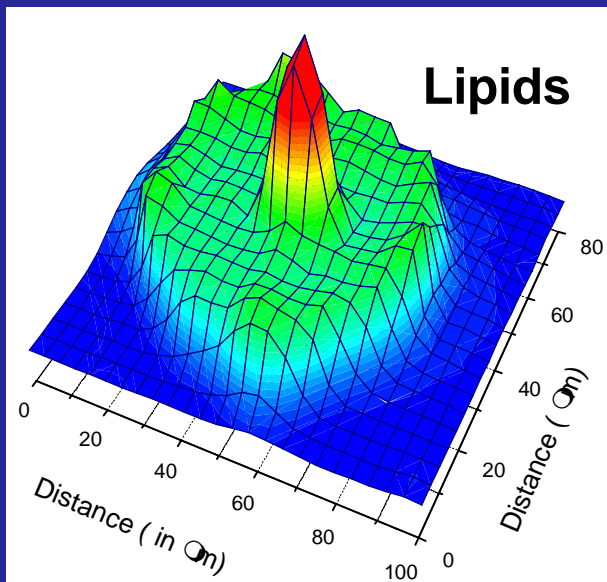
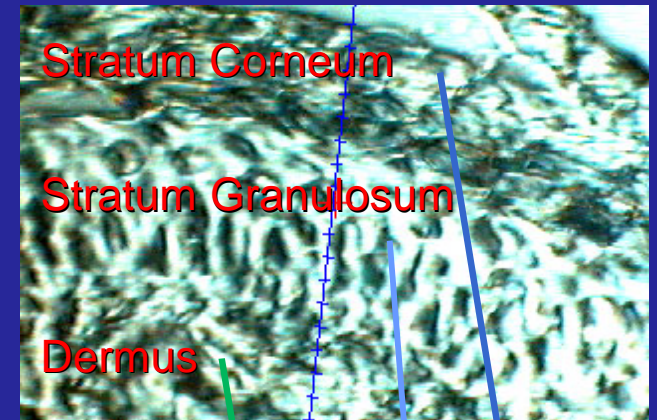
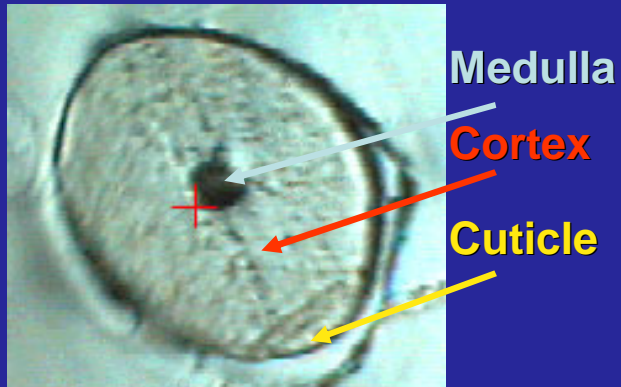
Lipids location



**Lipids concentrated at the contractil ring, where the cleavage furrow is located**



# Chemical Imaging of Hair and Skin



Paul Dumas et al., Physics World, May 2001, 2223

# Collaborators and Funding

## UWM

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