

Properties of Synchrotron Radiation

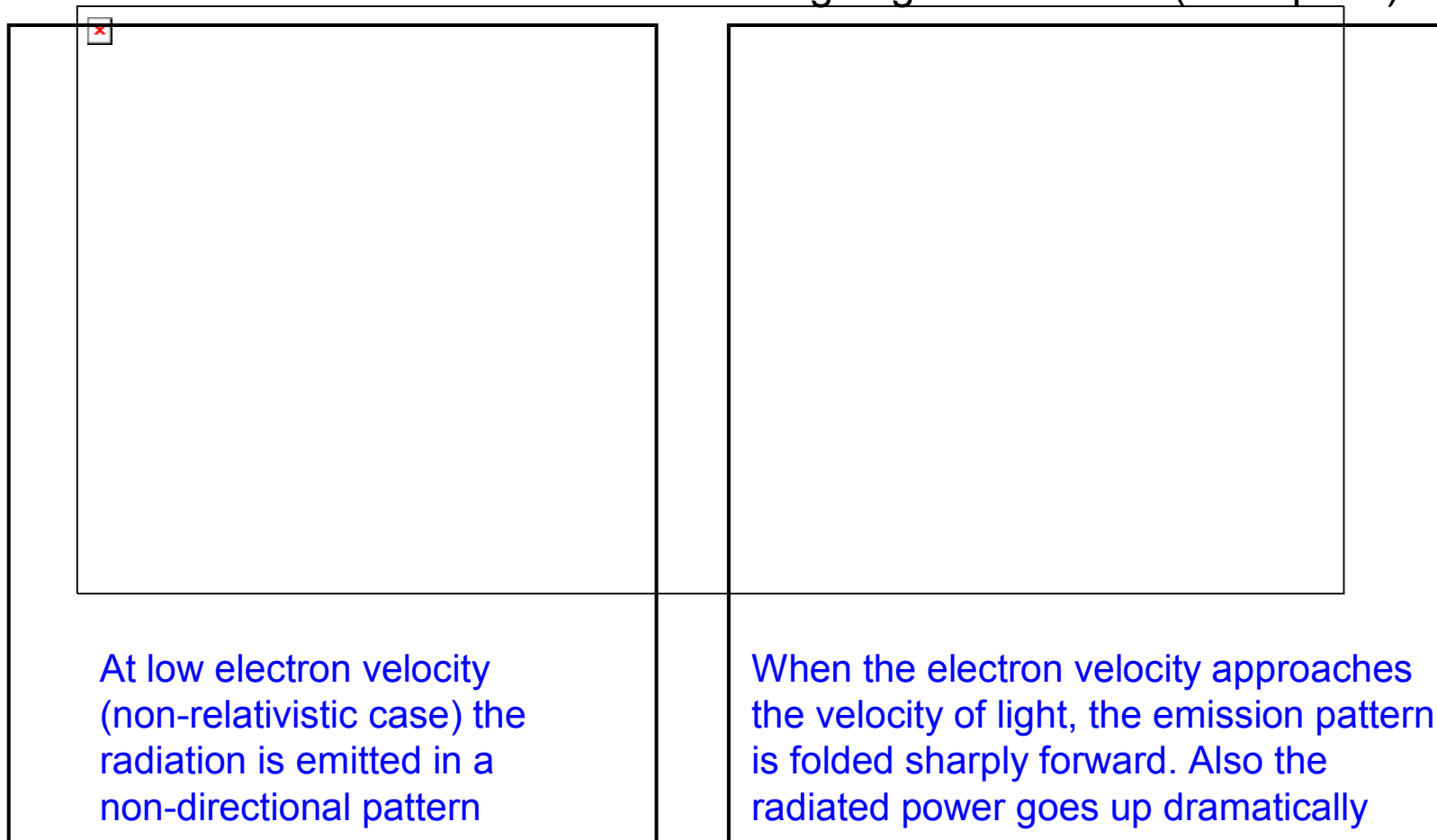
Presentation at JASS02 Seminar; Jordan, Oct. 19-28, 2002
Herman Winick, SSRL/SLAC, Stanford University

Comparison of Synchrotron Radiation from Synchrotrons and Storage Rings

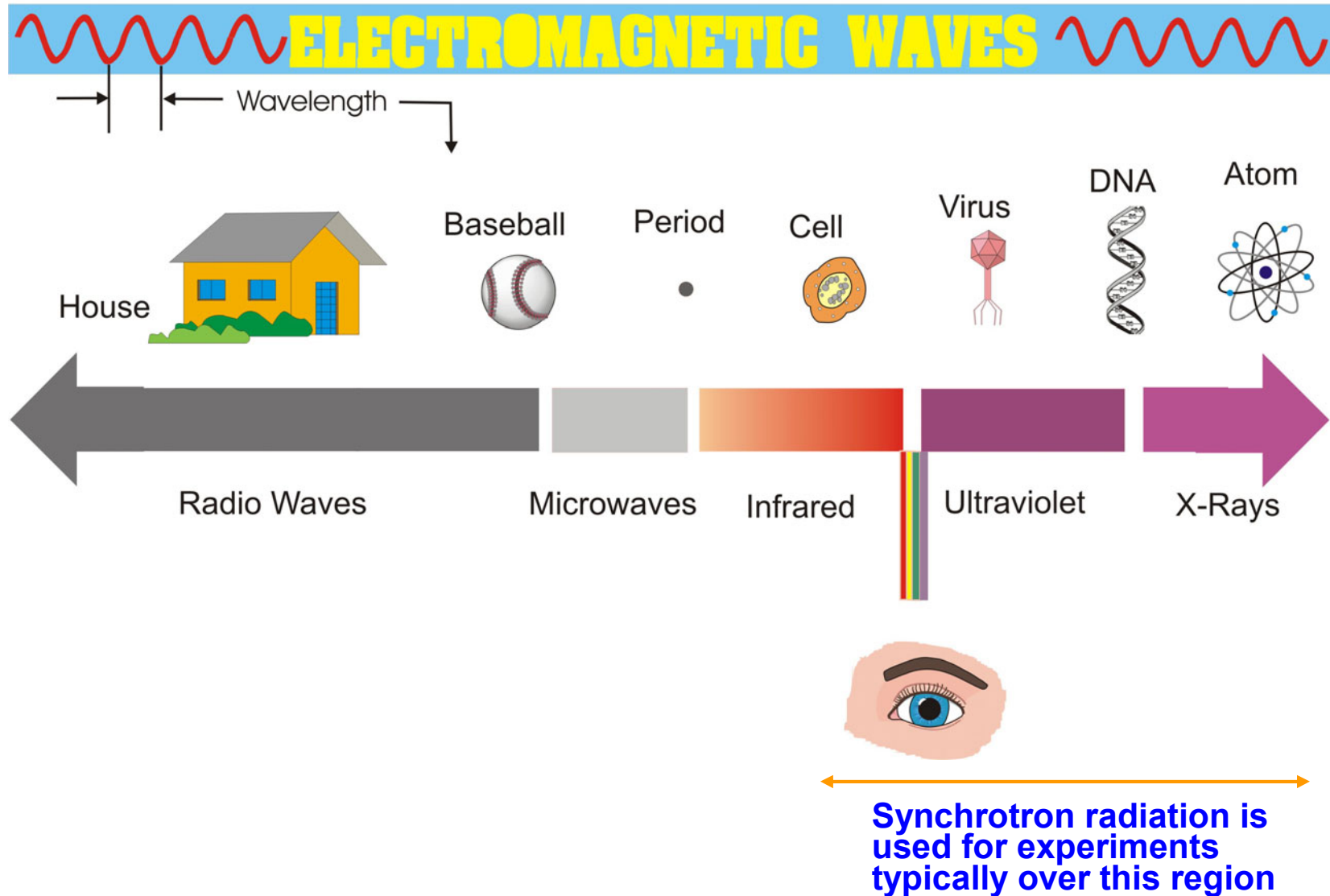
	<i>Synchrotron</i>	<i>Storage Ring</i>
Spectrum	Varies as e ⁻ energy changes on each cycle	Constant
Intensity	Varies as e ⁻ energy changes on each cycle, also; cycle to cycle variations	Decays slowly over many hours
Source Position	Varies during the acceleration cycle	Constant within 1-50 microns
High Energy Radiation Background	High – due to loss of all particles on each cycle	Low – particles are stored for many hours

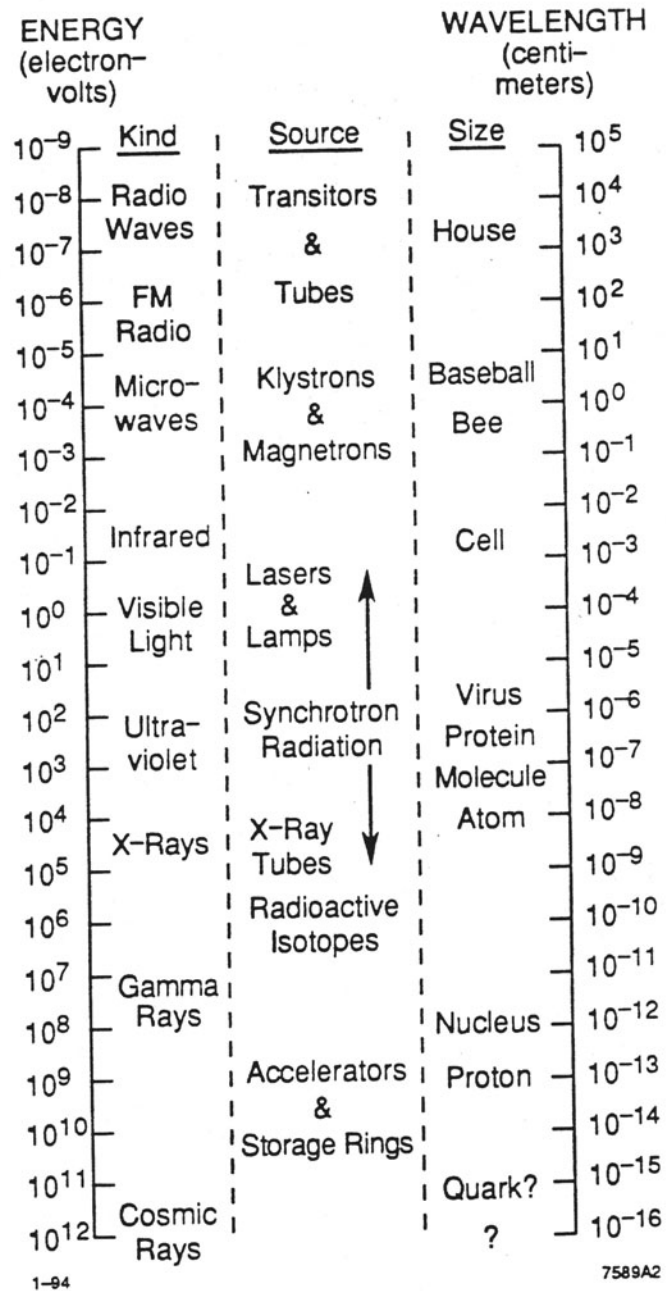
Radiation Fundamentals

- When electrons are accelerated (e.g. linear acceleration in a radio transmitter antenna) they emit electromagnetic radiation (*i.e.*, radio waves) in a rather non-directional pattern
- Electrons in circular motion are also undergoing acceleration (centripetal)



Electromagnetic Radiation - How It Relates to the World We Know





1-84

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The Electromagnetic spectrum showing the region occupied by synchrotron radiation

What Properties Make Synchrotron Radiation (SR) so Useful?

High brightness:

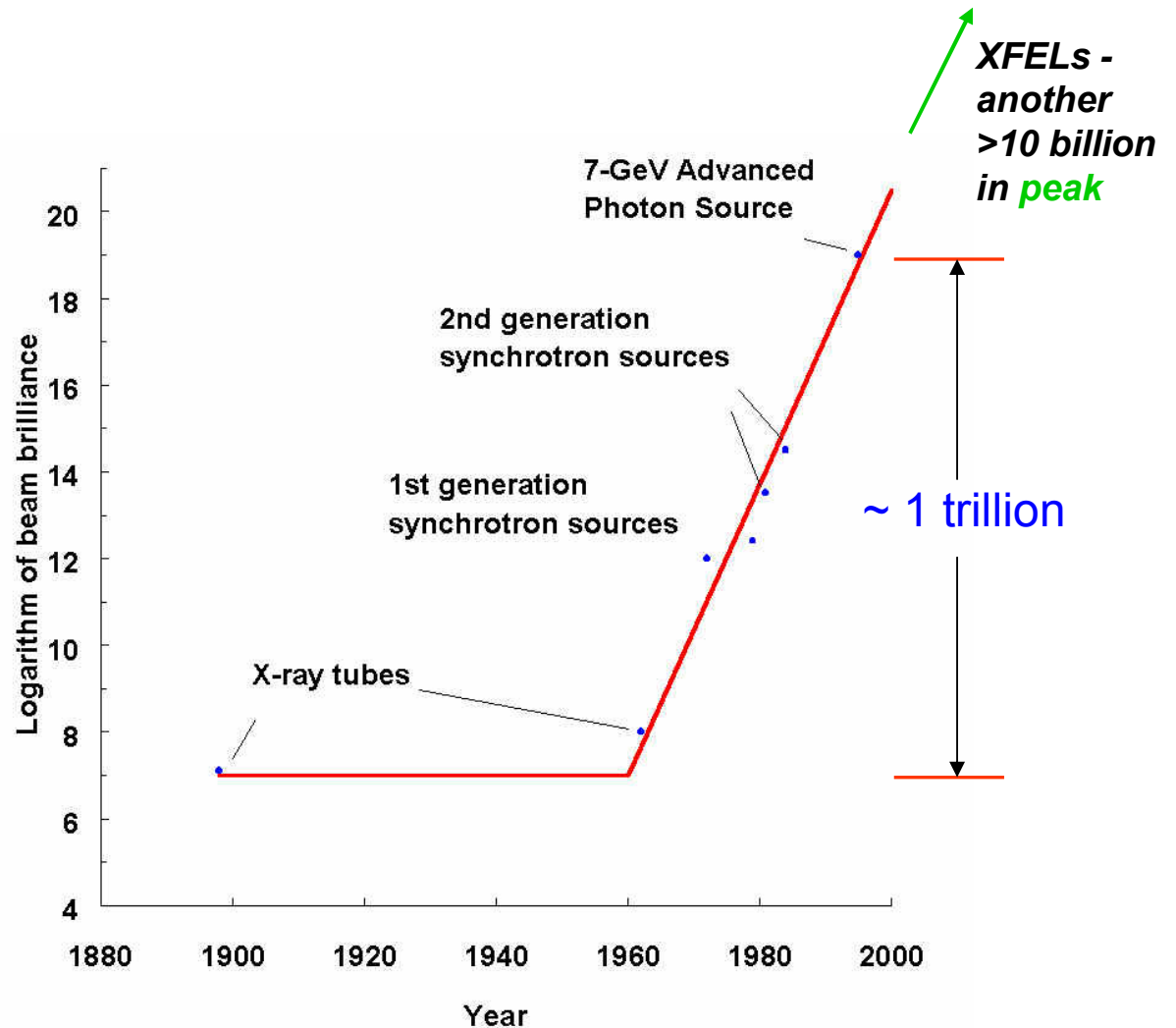
SR is extremely intense (hundreds of thousands of times higher than conventional X-ray tubes)

Wide energy spectrum:

SR is emitted with a wide range of energies

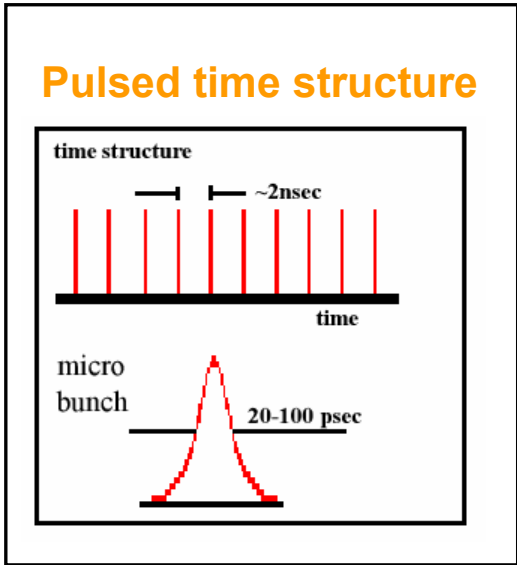
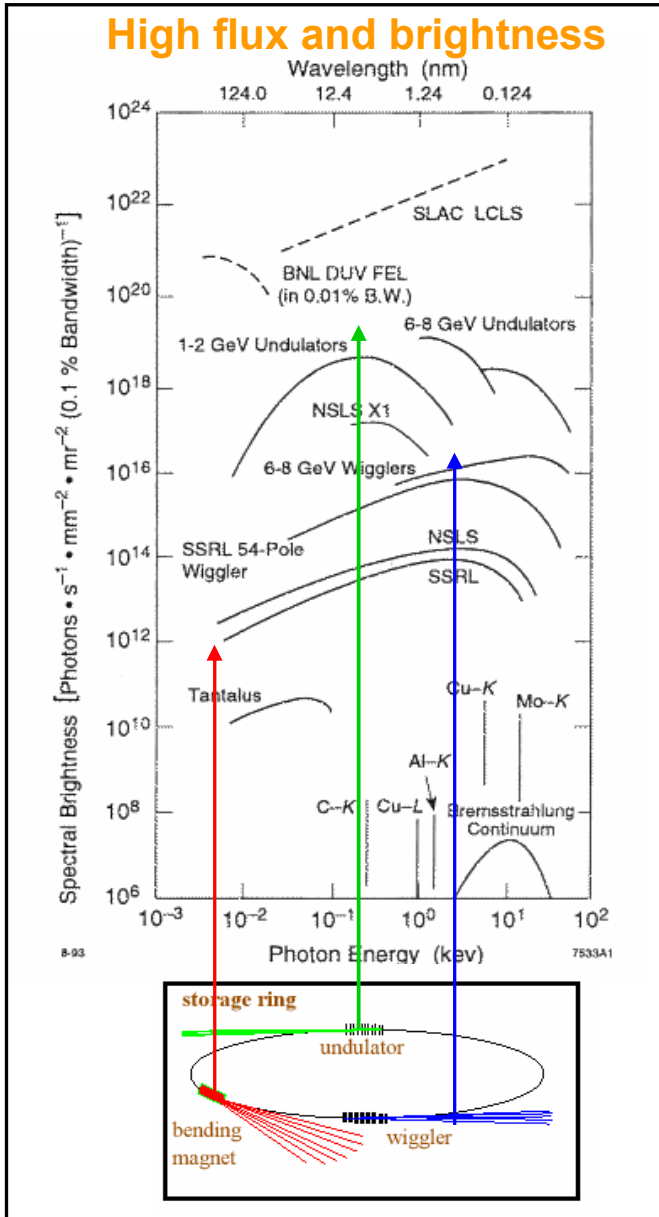
Highly polarized and short pulses:

SR is emitted in very short pulses, typically less than a nano-second (a billionth of a second)



SR offers many characteristics of visible lasers but into the x-ray regime!

Synchrotron Radiation - Basic Properties



- Broad spectral range**
- Polarized (linear, elliptical, circular)**
- Small source size**
- Partial coherence**
- High stability**

$$\text{Flux} = \frac{\# \text{ of photons in given } \Delta\lambda/\lambda}{\text{sec, mrad } \theta}$$

$$\text{Brightness} = \frac{\# \text{ of photons in given } \Delta\lambda/\lambda}{\text{sec, mrad } \theta, \text{ mrad } \varphi, \text{ mm}^2}$$

(a measure of concentration of the radiation)

SYNCHROTRON RADIATION

BASIC PROPERTIES

1. HIGH FLUX, BRIGHTNESS, STABILITY
2. BROAD SPECTRAL RANGE - Tunability
3. POLARIZATION (linear, elliptical, circular)
4. PULSED TIME STRUCTURE (0.01 - 1 nsec)
5. SMALL SOURCE SIZE (\leq mm)
6. PARTIAL COHERENCE
7. HIGH VACUUM ENVIRONMENT

Flux = No. of Photons at given λ within a given $\Delta\lambda/\lambda$
s, mrad Θ

Brightness = No. of Photons at given λ within a given $\Delta\lambda/\lambda$
s, mrad Θ , mrad ϕ , mm²
(a measure of the concentration of the radiation)

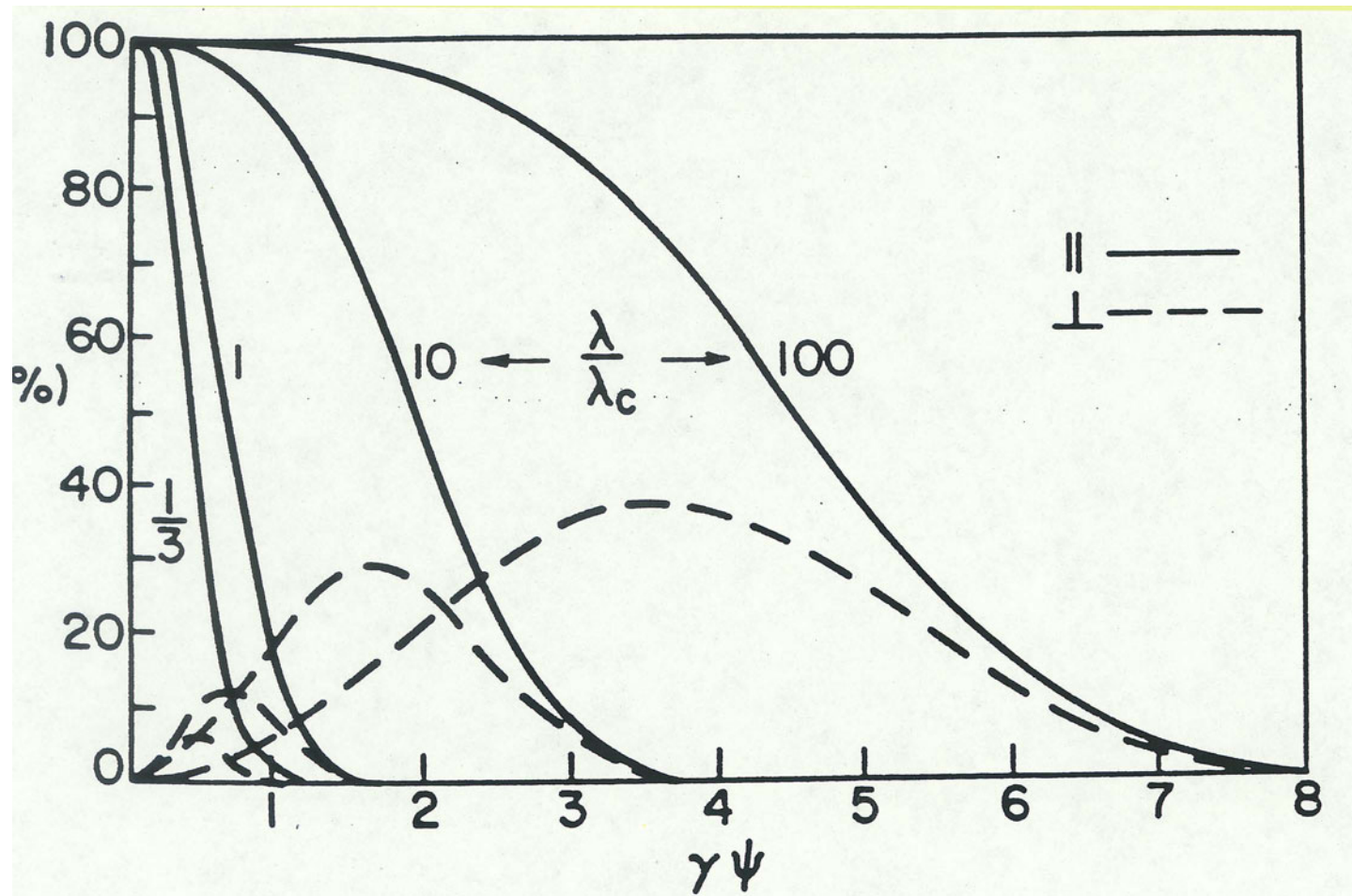
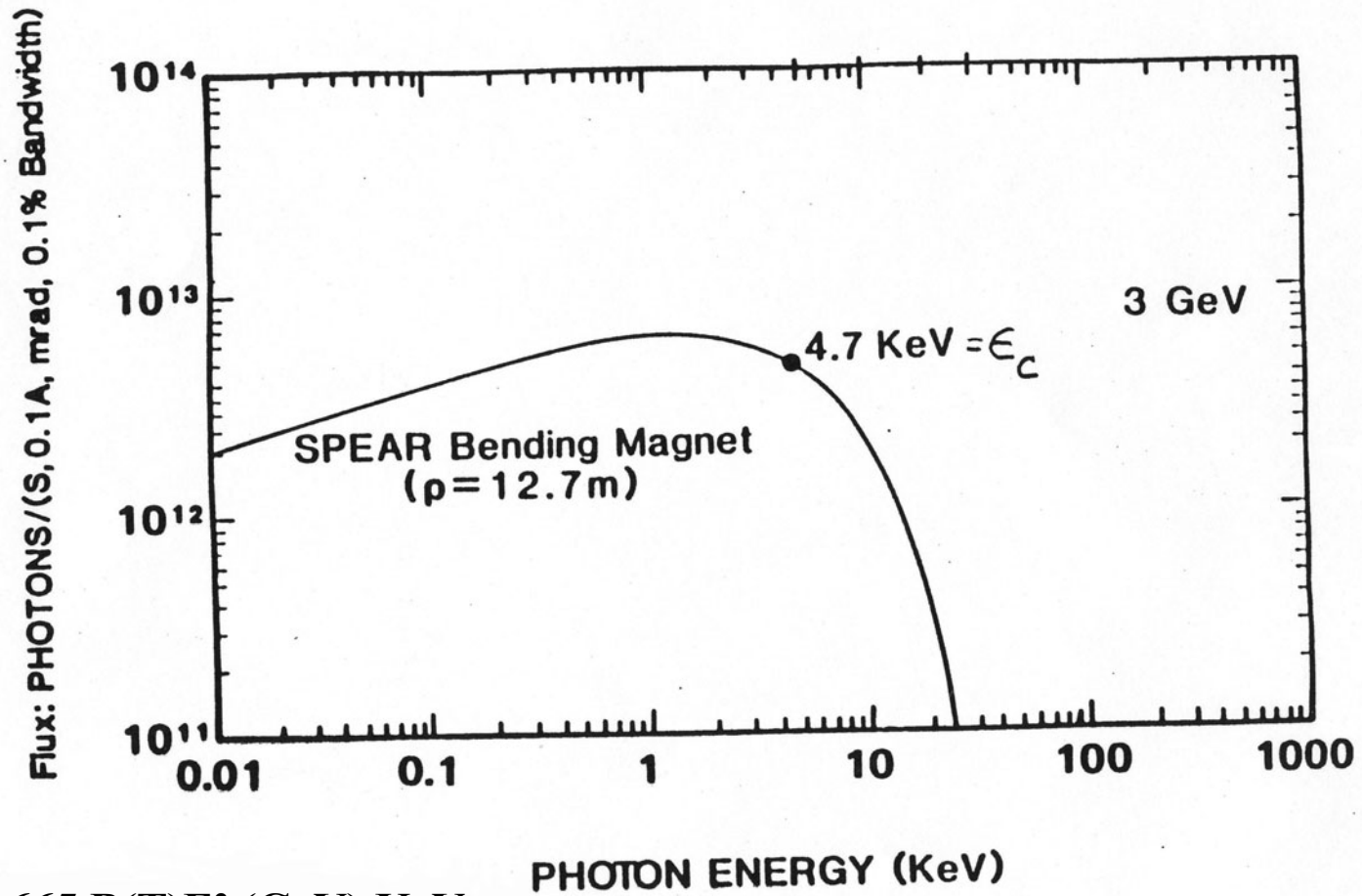
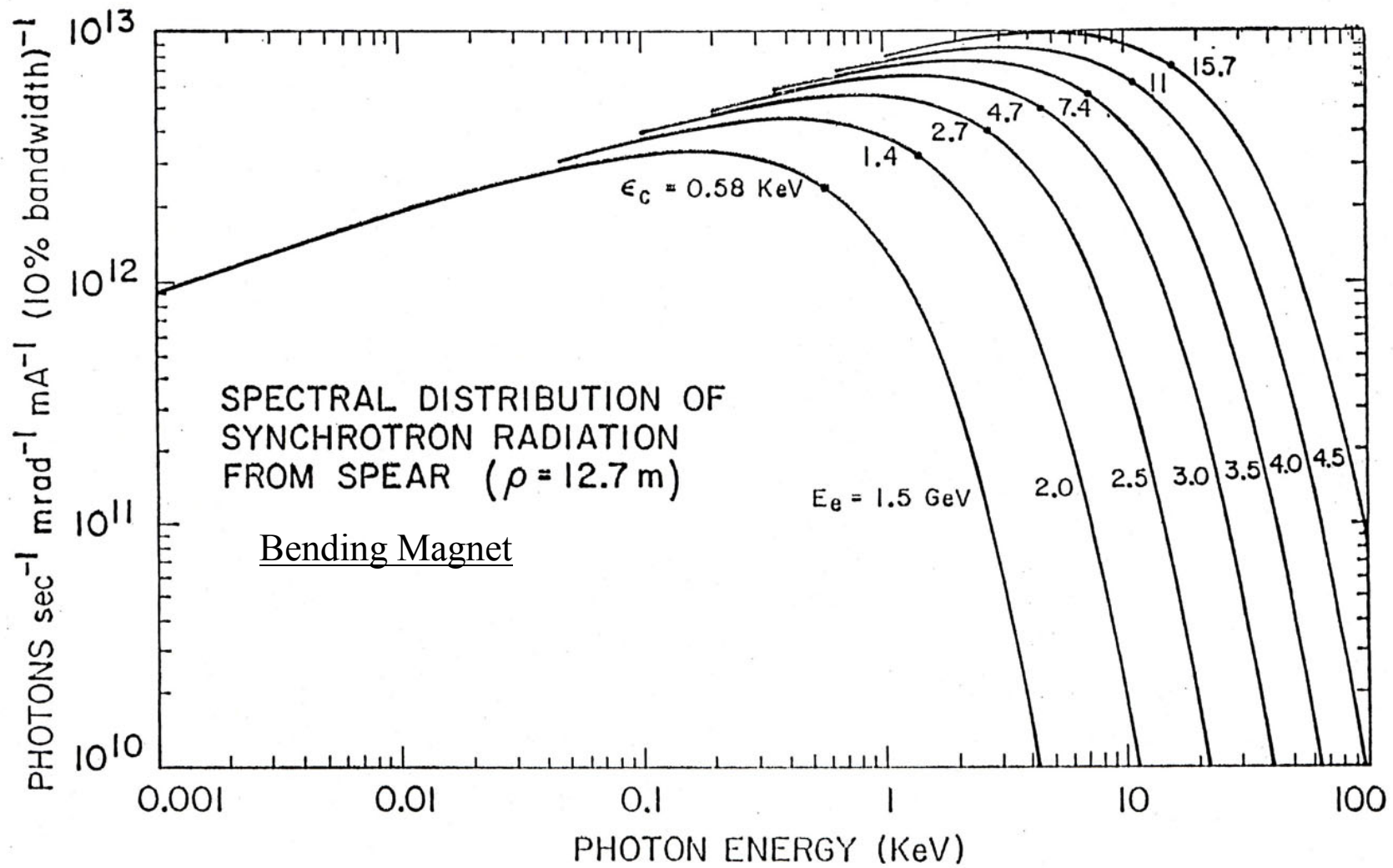


Figure 7. Vertical angular distribution of parallel and perpendicular polarization components.

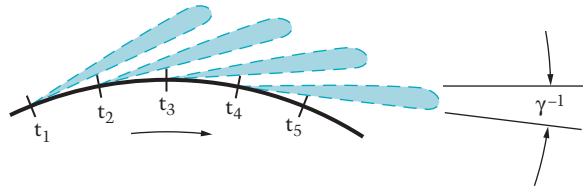


$$\epsilon = 0.665 B(T)E^2 (\text{GeV}) \text{ KeV}$$

$$= 2.2 E^3 (\text{GeV})/\rho(\text{m})$$



Bending Magnets and Insertion Devices on Storage Rings



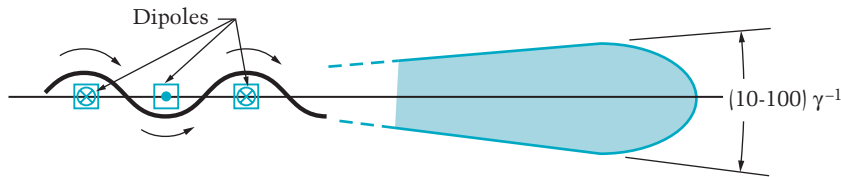
bending magnet - a "sweeping searchlight"

Continuous spectrum characterized by ε_c = critical energy

$$\varepsilon_c(\text{keV}) = 0.665 B(\text{T})E^2(\text{GeV})$$

eg: for $B = 1.35\text{T}$ $E = 2\text{GeV}$

$$\varepsilon_c = 3.6\text{keV}$$



wiggler - incoherent superposition

Quasi-monochromatic spectrum with peaks at lower energy than a wiggler

$$\lambda_1 = \frac{\lambda_u}{2\gamma^2} \left(1 + \frac{K^2}{2}\right) \sim \frac{\lambda_u}{\gamma^2} \text{ (fundamental)}$$

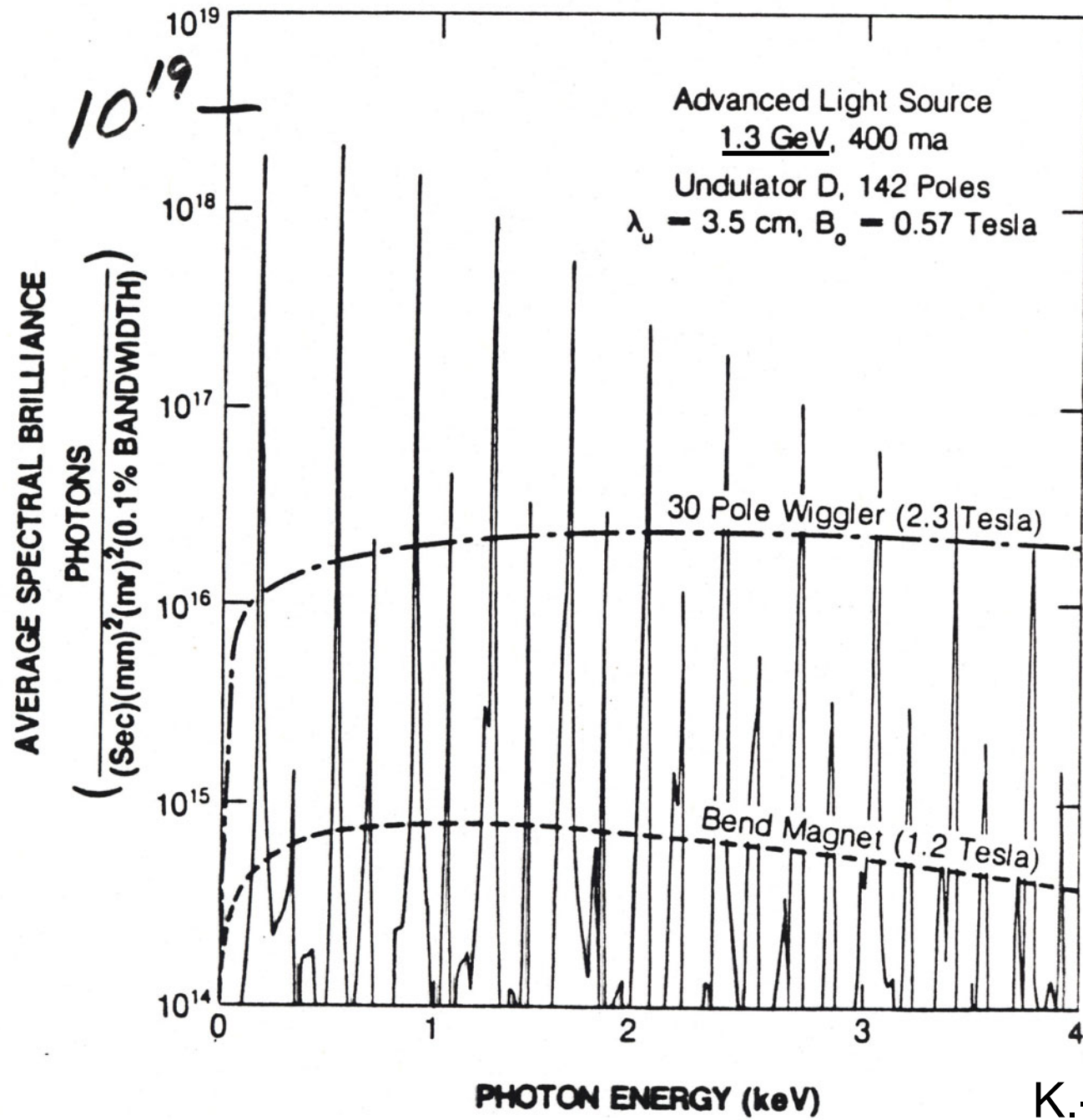
+ harmonics at higher energy

$$\varepsilon_1(\text{keV}) = \frac{0.95 E^2(\text{GeV})}{\lambda_u(\text{cm}) \left(1 + \frac{K^2}{2}\right)}$$

$K = \gamma\theta$ where θ is the angle in each pole



undulator - coherent interference



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End of this part of presentation